



Introduction to gyrokinetics - a quick and dirty derivation

Josefine H.E. Proll



EUROfusion



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Gradients in plasma lead to unwanted turbulence

Global Gyrokinetic Simulation of Turbulence in **ASDEX Upgrade**



GENE

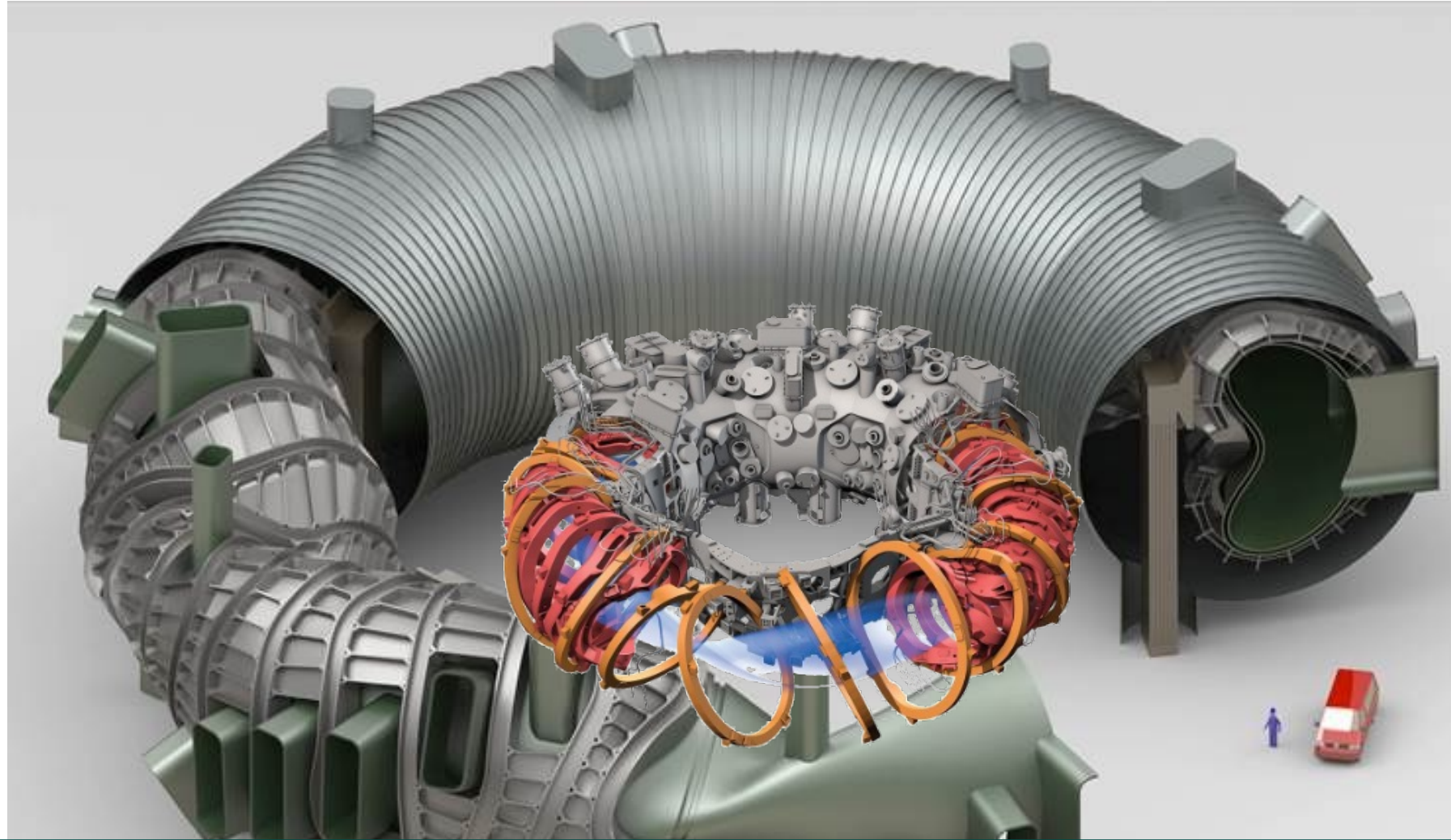
gene.rzg.mpg.de

Density and temperature
gradients:

Turbulence!

Wendelstein 7-X:
1Bn EUR

**Reactor-sized:
20 Bn EUR**



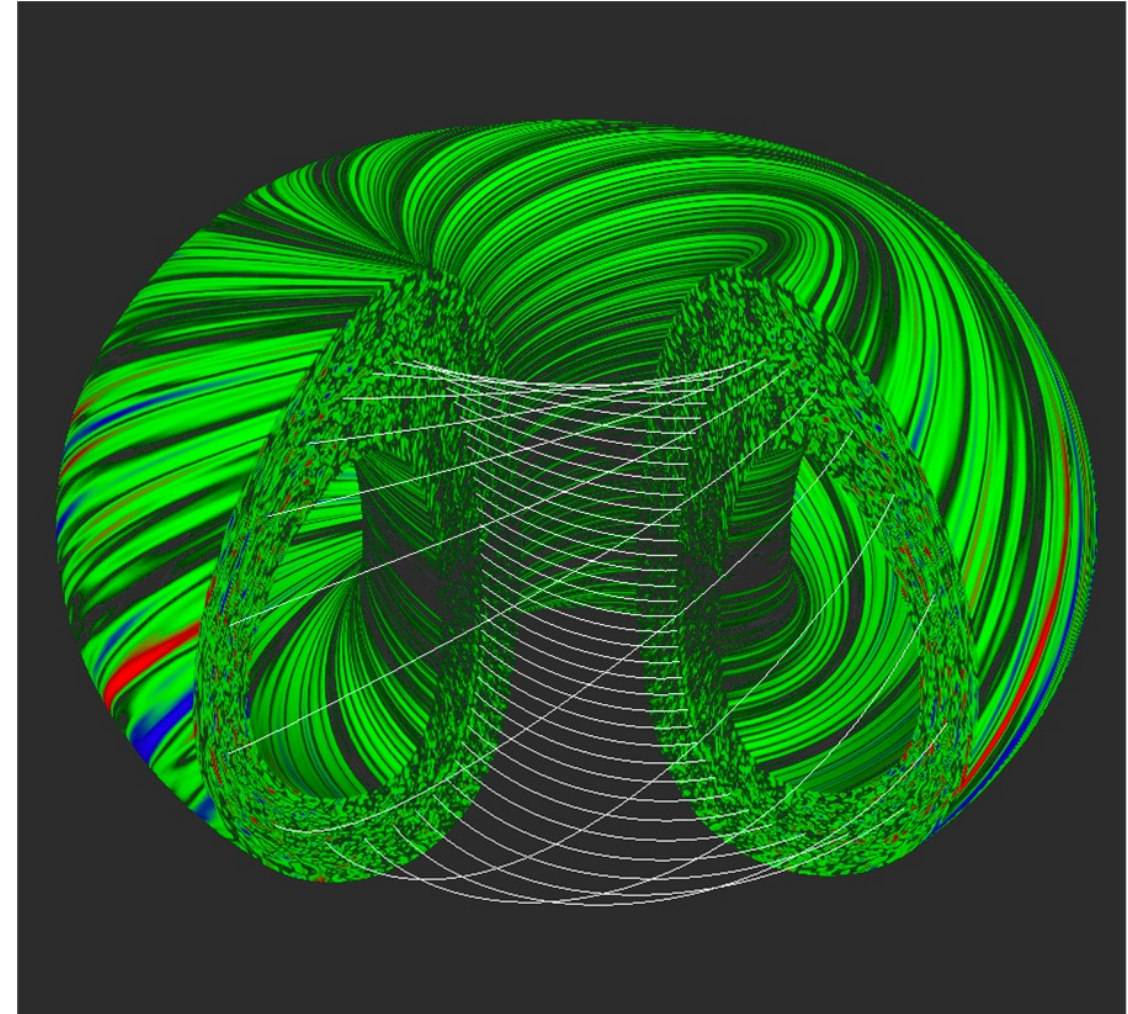
Turbulence as bottleneck – reduce turbulence – build smarter, not larger



Fluctuations in fusion plasmas

- Observed features:
- Everything fluctuates
(n, Te, Ti, electromagnetic fields)
- Fluctuation levels tend to be small – $\frac{\tilde{n}}{n} \approx 0.1 \dots 1\%$
- Fluctuation frequencies: 10 – 1000s kHz
are smaller than the gyrofrequencies
- Perpendicular fluctuation scales much smaller than
the system size – on the order of the gyroradius
- Fluctuations extended along the field lines

Can't use MHD,
but allows the use of gyrokinetics!

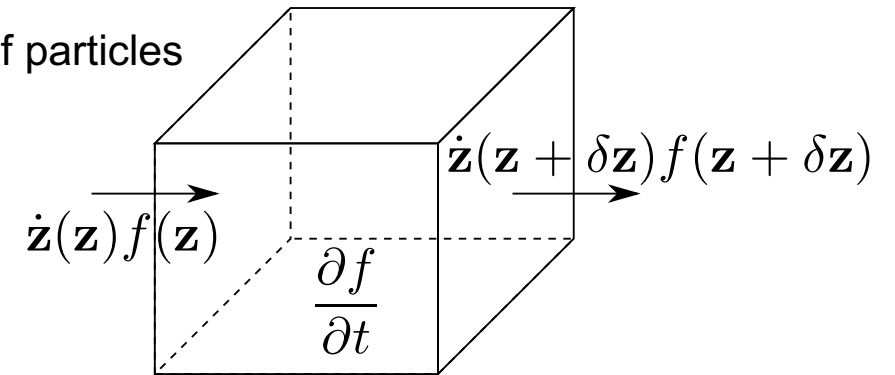




The kinetic equation – a continuity equation of the distribution in 6D phase space

- Kinetic description – intermediate step between fluid description and resolving single-particle dynamics
- Describe evolution of particles in terms of distribution function f_a of particles in 6D phase space \mathbf{z} (3D real space \mathbf{x} , 3D velocity space \mathbf{v})

$$\frac{\partial f_a}{\partial t} + \nabla_{\mathbf{z}} (\dot{\mathbf{z}} f_a) = C_a(f_a), \quad \dot{\mathbf{z}} = (\dot{\mathbf{x}}, \dot{\mathbf{v}})$$



$$\dot{\mathbf{v}} = \frac{e_a}{m_a} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \quad \nabla_{\mathbf{x}} \cdot \mathbf{v} = 0 \quad \nabla_{\mathbf{v}} \cdot (\mathbf{E} + \mathbf{v} \times \mathbf{B}) = 0$$

$$\frac{\partial f_a}{\partial t} + \mathbf{v} \cdot \nabla_{\mathbf{x}} f_a + \frac{e_a}{m_a} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_{\mathbf{v}} f_a = C_a(f_a)$$



The kinetic equation... now in more convenient coordinates

$$\frac{\partial f_a}{\partial t} + \mathbf{v} \cdot \nabla_{\mathbf{x}} f_a + \frac{e_a}{m_a} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_{\mathbf{v}} f_a = C_a(f_a)$$

- Freedom in choosing suitable coordinates for \mathbf{x} and \mathbf{v}
- Make use of coordinates related to the gyromotion:

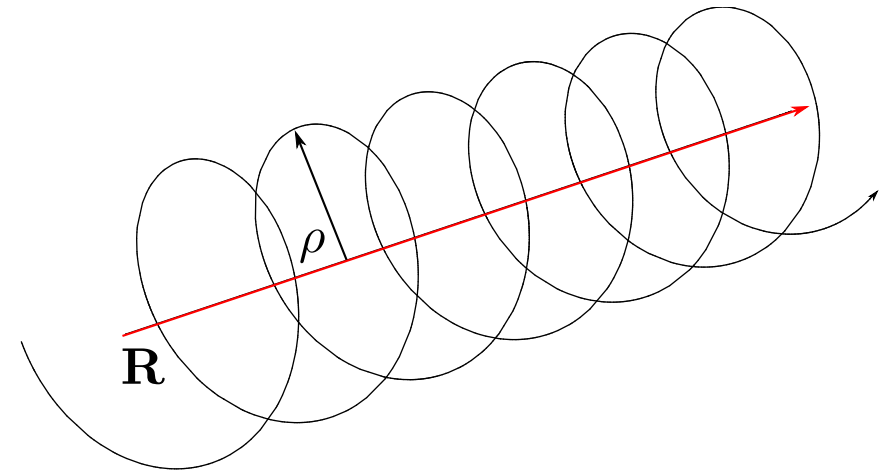
- Gyro angle ϑ

- Energy $\mathcal{E} = \frac{m_a v^2}{2} + e_a \phi$

- Magnetic moment $\mu = \frac{m_a v_{\perp}^2}{2B}$

- Gyro-centre $\mathbf{R} = \mathbf{r} + \frac{\hat{\mathbf{b}} \times \mathbf{v}}{\Omega_a}$

$$\frac{\partial f_a}{\partial t} + \dot{\mathbf{R}} \cdot \frac{\partial f_a}{\partial \mathbf{R}} + \dot{\vartheta} \frac{\partial f_a}{\partial \vartheta} + \dot{\mathcal{E}} \frac{\partial f_a}{\partial \mathcal{E}} + \dot{\mu} \frac{\partial f_a}{\partial \mu} = C_a$$





The kinetic equation in more convenient coordinates cont'd

$$\frac{\partial f_a}{\partial t} + \dot{\mathbf{R}} \cdot \frac{\partial f_a}{\partial \mathbf{R}} + \dot{\vartheta} \frac{\partial f_a}{\partial \vartheta} + \dot{\mathcal{E}} \frac{\partial f_a}{\partial \mathcal{E}} + \dot{\mu} \frac{\partial f_a}{\partial \mu} = C_a$$

We can simplify:

$$m_a \dot{\mathbf{v}} = e_a (\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

$$\dot{\vartheta} \simeq -\Omega_a$$

$$\dot{\mathcal{E}} = \frac{d}{dt} \left(\frac{m_a v^2}{2} \right) + e_a \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} + e_a \frac{\partial \phi}{\partial \mathbf{r}} \cdot \frac{d\mathbf{r}}{dt}$$

$$= m_a \dot{\mathbf{v}} \cdot \mathbf{v} + e_a \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} + e_a \mathbf{v} \cdot \nabla \phi$$

$$= -e_a \mathbf{v} \cdot \nabla \phi + e_a \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} + e_a \mathbf{v} \cdot \nabla \phi$$

$$= e_a \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}}$$



Simplifying the kinetic equation some more

We decompose the distribution function $f_a = f_{a0} + g_a$,

f_{a0} equilibrium function including adiabatic responses to small electric fields $\propto -e_a\phi/T_a$,

g_a Small non-adiabatic part with $g_a \ll f_{a0}$

Assumptions:

- Equilibrium varies slowly in time c.f. perturbed part $\frac{\partial g_a}{\partial t} \gtrsim \frac{\partial f_{a0}}{\partial t}$
- But allow $\nabla g_a \sim \nabla f_{a0}$.
- And further: small perturbations in electrostatic field, equilibrium length scales L large compared with gyroradius, frequencies of instabilities ω small compared with gyrofrequency:

$$\frac{e_a\phi}{T_a} \sim \frac{\rho}{L} \sim \frac{\omega}{\Omega_a} \sim \delta \ll 1$$



Your turn: ordering of the terms in δ

$$\frac{\partial f_a}{\partial t} + \dot{\mathbf{R}} \cdot \frac{\partial f_a}{\partial \mathbf{R}} + \dot{\vartheta} \frac{\partial f_a}{\partial \vartheta} + \dot{\mathcal{E}} \frac{\partial f_a}{\partial \mathcal{E}} + \dot{\mu} \frac{\partial f_a}{\partial \mu} = C_a$$

$$\dot{\vartheta} \simeq -\Omega_a \quad \dot{\mathcal{E}} = e_a \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}}$$

$$f_a = f_{a0} + g_a, \quad g_a \ll f_{a0} \quad \nabla g_a \sim \nabla f_{a0}, \quad \frac{\partial g_a}{\partial t} \gtrsim \frac{\partial f_{a0}}{\partial t}$$

$$\frac{\partial g_a}{\partial \mathcal{E}} \ll \frac{\partial f_{a0}}{\partial \mathcal{E}}$$

$$\frac{\partial g_a}{\partial \mu} \ll \frac{\partial f_{a0}}{\partial \mu}$$



In lowest orders

Equilibrium is independent of gyroangle: $\partial f_{a0} / \partial \vartheta = 0$

In next order
$$\frac{\partial f_{a0}}{\partial t} + \frac{\partial g_a}{\partial t} + \dot{\mathbf{R}} \cdot \frac{\partial}{\partial \mathbf{R}} (f_{a0} + g_a) - \Omega_a \frac{\partial g_a}{\partial \vartheta} + e \frac{\partial \phi}{\partial t} \frac{\partial f_{a0}}{\partial \mathcal{E}} + \dot{\mu} \frac{\partial f_{a0}}{\partial \mu} = C_a,$$



Now choosing the equilibrium to be represented by a Maxwellian

(follows from drift kinetics in lowest order, at high collisionality)

$$f_{a0} = n_a(\psi) \left(\frac{m_a}{2\pi T_a(\psi)} \right)^{3/2} e^{-\mathcal{E}/T_a(\psi)}$$
$$\simeq n_a(\psi) \left(\frac{m_a}{2\pi T_a(\psi)} \right)^{3/2} e^{-m_a v^2 / 2T_a(\psi)} \left(1 - \frac{e_a \Phi}{T_a} \right)$$

- Note:
density n and temperature T are assumed to be flux quantities with flux label ψ

- Then:

$$\frac{\partial f_{a0}}{\partial t} = 0$$

$$\frac{\partial f_{a0}}{\partial \mu} = 0$$

$$\frac{\partial f_{a0}}{\partial \mathcal{E}} = -\frac{f_{a0}}{T_a}$$



Splitting the movement into parallel and gyro motion

Assuming the particle mostly moves along the field $\dot{\mathbf{R}} = v_{\parallel} \hat{\mathbf{b}} + \mathcal{O}(\delta v_T)$

We obtain
$$\frac{\partial g_a}{\partial t} + \dot{\mathbf{R}} \cdot \frac{\partial}{\partial \mathbf{R}} (f_{a0} + g_a) - \Omega \frac{\partial g_a}{\partial \theta} - \frac{e}{T} \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} f_{a0} = C_a.$$

All terms are of order $\mathcal{O}(\delta \omega f_{a0})$ with $\omega \sim v_T/L$, except $\Omega \frac{\partial g_a}{\partial \theta} \sim \omega f_{a0}$ as largest term

We can thus expand $g_a = g_{a0} + g_{a1} + \dots$ yielding in lowest order $\Omega \frac{\partial g_{a0}}{\partial \theta} = 0$

In next order after gyroaveraging hile keeping the gyro centre constant

$$\frac{\partial g_{a0}}{\partial t} + \langle \dot{\mathbf{R}} \rangle_{\mathbf{R}} \cdot \frac{\partial}{\partial \mathbf{R}} (f_{a0} + g_{a0}) - \frac{e_a}{T_a} \left\langle \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} \right\rangle_{\mathbf{R}} f_{a0} = \langle C_a \rangle_{\mathbf{R}}.$$



Averaging out the fast gyromotion

- Magnetic field \rightarrow reduction to 5D phase space

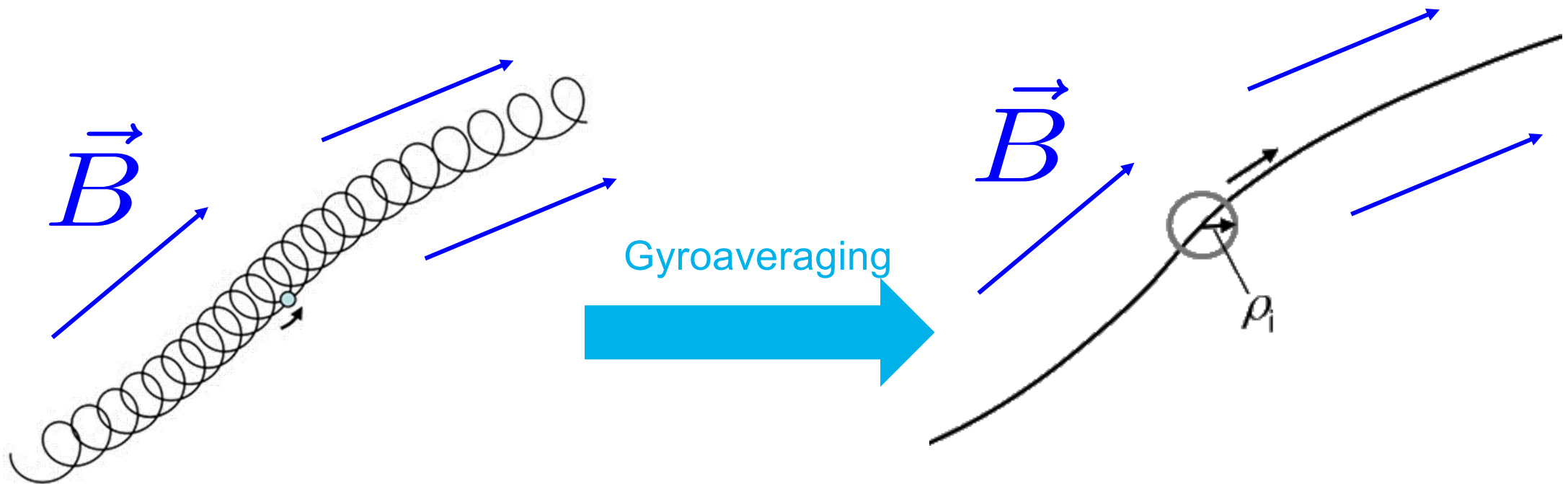


Image source: Garbet *et al. Nucl. Fusion* (2010).



Splitting the velocity

Splitting the velocity: $\langle \dot{\mathbf{R}} \rangle_{\mathbf{R}} = \frac{1}{2\pi} \oint \dot{\mathbf{R}} d\vartheta = v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_E + \mathbf{v}_d$

$$\mathbf{v}_E = \frac{\hat{\mathbf{b}} \times \nabla \langle \phi \rangle_{\mathbf{R}}}{B}$$

$$\mathbf{v}_{da} = \frac{\hat{\mathbf{b}}}{\Omega_a} \times \left(\frac{v_{\perp}^2}{2} \nabla \ln B + v_{\parallel}^2 \boldsymbol{\kappa} \right)$$

Using that f_{a0} is a flux function and thus does not vary along the field lines $\mathbf{b} \cdot \nabla f_{a0} = 0$,

$$\frac{\partial g_{a0}}{\partial t} + \left(v_{\parallel} \hat{\mathbf{b}} + \underbrace{\mathbf{v}_E + \mathbf{v}_d}_A \right) \cdot \nabla g_{a0} - \langle C \rangle_{\mathbf{R}} = - \left(\mathbf{v}_E + \underbrace{\mathbf{v}_d}_B \right) \cdot \nabla f_{a0} + \frac{e}{T} \left\langle \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} \right\rangle_{\mathbf{R}} f_{a0}$$

nonlinear
Neoclassical response



Executing the gyroaverages: $\frac{e}{T} \left\langle \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} \right\rangle_{\mathbf{R}} f_{a0}$

Have slow variation along the field but fast across it, can write: $\phi(\mathbf{r}, t) = \hat{\phi}(\mathbf{r}, \omega) \underbrace{e^{i(S(\mathbf{r})/\delta - \omega t)}}_{\text{Fast variation}}$

Define $\mathbf{k}_{\perp} = \frac{\nabla S}{\delta}$ with $k_{\perp} \rho = \mathcal{O}(1)$

Ballooning transform allows assumption of $\nabla_{\parallel} S = 0$.

$$\text{Then } \left\langle \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} \right\rangle_{\mathbf{R}} = -i\omega \langle \phi(\mathbf{R} + \boldsymbol{\rho}) \rangle_{\mathbf{R}} \simeq -i\omega \hat{\phi}(\mathbf{R}) e^{-i\omega t} \left\langle e^{iS(\mathbf{R} + \boldsymbol{\rho})/\delta} \right\rangle_{\mathbf{R}}$$

$$\text{where } \left\langle e^{iS(\mathbf{R} + \boldsymbol{\rho})/\delta} \right\rangle_{\mathbf{R}} \simeq e^{iS(\mathbf{R}/\delta)} \left\langle e^{i\mathbf{k}_{\perp} \cdot \boldsymbol{\rho}} \right\rangle_{\mathbf{R}} = J_0 \left(\frac{k_{\perp} v_{\perp}}{\Omega_a} \right) e^{iS(\mathbf{R}/\delta)}$$

$$\text{So that } \left\langle \left(\frac{\partial \phi}{\partial t} \right)_{\mathbf{r}} \right\rangle_{\mathbf{R}} = -i\omega J_0 \left(\frac{k_{\perp} v_{\perp}}{\Omega_a} \right) \phi(\mathbf{R}, t)$$

Bessel function of zeroth order

$$\int_0^{2\pi} e^{ix \sin \vartheta} d\vartheta = \int_0^{2\pi} \cos(x \sin \vartheta) d\vartheta = 2\pi J_0(x)$$



Executing the gyroaverages

$$-\mathbf{v}_E \cdot \nabla f_{a0} = -\frac{\hat{\mathbf{b}} \times \nabla \langle \phi \rangle_{\mathbf{R}}}{B} \cdot \nabla f_{a0} = \frac{\nabla \langle \phi \rangle_{\mathbf{R}}}{B} \cdot \hat{\mathbf{b}} \times \nabla f_{a0}.$$

$$\begin{aligned} \nabla \langle \phi \rangle_{\mathbf{R}} &= \nabla \langle \phi(\mathbf{R} + \boldsymbol{\rho}) \rangle_{\mathbf{R}} \\ &\simeq \nabla \left[\hat{\phi}(\mathbf{R}) \left\langle e^{iS(\mathbf{R}+\boldsymbol{\rho})/\delta} \right\rangle_{\mathbf{R}} e^{-i\omega t} \right] \\ &\simeq \nabla \left(\hat{\phi}(\mathbf{R}) e^{-i\omega t} e^{iS(\mathbf{R})/\delta} J_0 \left(\frac{k_{\perp} v_{\perp}}{\Omega_a} \right) \right) \\ &\simeq \underbrace{\hat{\phi}(\mathbf{R}) e^{-i\omega t} e^{iS(\mathbf{R})/\delta}}_{\phi(\mathbf{R}, t)} i \frac{\nabla S}{\delta} J_0 \left(\frac{k_{\perp} v_{\perp}}{\Omega_a} \right) \quad \text{because } \frac{\nabla S}{\delta} \gg \frac{\nabla \hat{\phi}(\mathbf{R})}{\hat{\phi}(\mathbf{R})} \\ &\simeq i \mathbf{k}_{\perp} J_0 \left(\frac{k_{\perp} v_{\perp}}{\Omega_a} \right) \phi(\mathbf{R}, t). \end{aligned}$$

We thus obtain

$$-\mathbf{v}_E \cdot \nabla f_{a0} = i J_0 \left(\frac{k_{\perp} v_{\perp}}{\Omega} \right) \phi(\mathbf{R}, t) \frac{1}{B} \mathbf{k}_{\perp} \cdot \hat{\mathbf{b}} \times \nabla f_{a0}$$



Simplify further with gradients in the distribution function

With f_{a0} being a flux function: $\nabla f_{a0} = \frac{\partial f_{a0}}{\partial \psi} \nabla \psi$

Magnetic field in Clebsch coordinates: $\mathbf{B} = \nabla \psi \times \nabla \alpha \quad \rightarrow \quad \mathbf{k}_\perp = k_\psi \nabla \psi + k_\alpha \nabla \alpha$

Obtain
$$\begin{aligned} -\mathbf{v}_E \cdot \nabla f_{a0} &= iJ_0 \phi \frac{\hat{\mathbf{b}} \cdot (\nabla \psi \times \mathbf{k}_\perp)}{B} \frac{\partial f_{a0}}{\partial \psi} \\ &= iJ_0 \phi k_\alpha \left[\frac{d \ln n_a}{d \psi} + \left(\frac{\mathcal{E}}{T_a} - \frac{3}{2} \right) \frac{d \ln T_a}{d \psi} \right] f_{a0} \\ &= iJ_0 \frac{e_a \phi}{T_a} \omega_{*a}^T f_{a0} \end{aligned}$$

With
$$\omega_{*a}^T = \omega_{*a} \left[1 + \eta_a \left(\frac{\mathcal{E}}{T_a} - \frac{3}{2} \right) \right] \quad \omega_{*a} = \frac{T_a k_\alpha}{e_a} \frac{d \ln n_a}{d \psi} \quad \eta_a = \frac{d \ln T_a}{d \psi} \bigg/ \frac{d \ln n_a}{d \psi}$$



Final step: same separation also in g_{a0}

With $g_{a0}(\mathbf{R}, \mathcal{E}, \mu, t) = \hat{g}_a(\mathbf{R}, \mathcal{E}, \mu) e^{i(S(\mathbf{R})/\delta - \omega t)}$

Obtain $\frac{\partial g_{a0}}{\partial t} + (v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_{da}) \cdot \nabla g_{a0} \simeq \left[v_{\parallel} \nabla_{\parallel} \hat{g}_a - i(\omega - \omega_{da}) \hat{g}_a \right] e^{i(S(\mathbf{R})/\delta - \omega t)}$

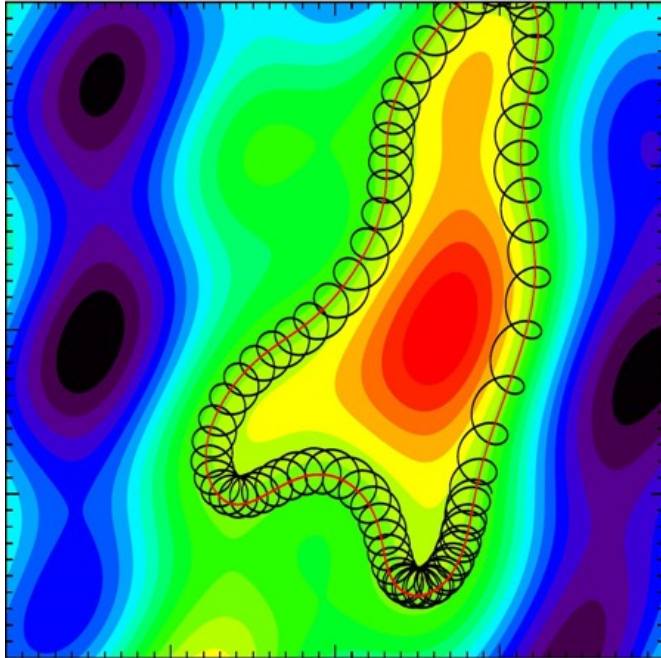
With $\omega_{da} = \mathbf{k}_{\perp} \cdot \mathbf{v}_{da}$.

Leaving out terms A and B (nonlinearity and neoclassical response) and collisions

$$v_{\parallel} \nabla_{\parallel} g_a - i(\omega - \omega_{da}) g_a = -\frac{ie_a}{T_a} J_0 \phi \left(\omega - \omega_{*a}^T \right) f_{a0}$$



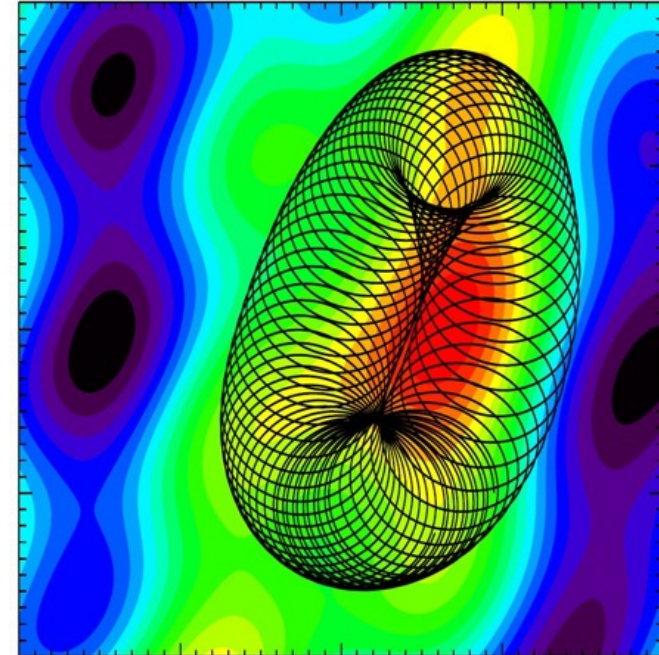
Suitability of gyroaveraging: small gyration



Gyromotion small compared with background: gyroaveraging allowed

Important condition:

$$\rho_s \ll L$$



Gyromotion large compared with background: gyroaveraging not allowed