

# Low Temperature Plasma IV: Applications

Yevgeny Raitses

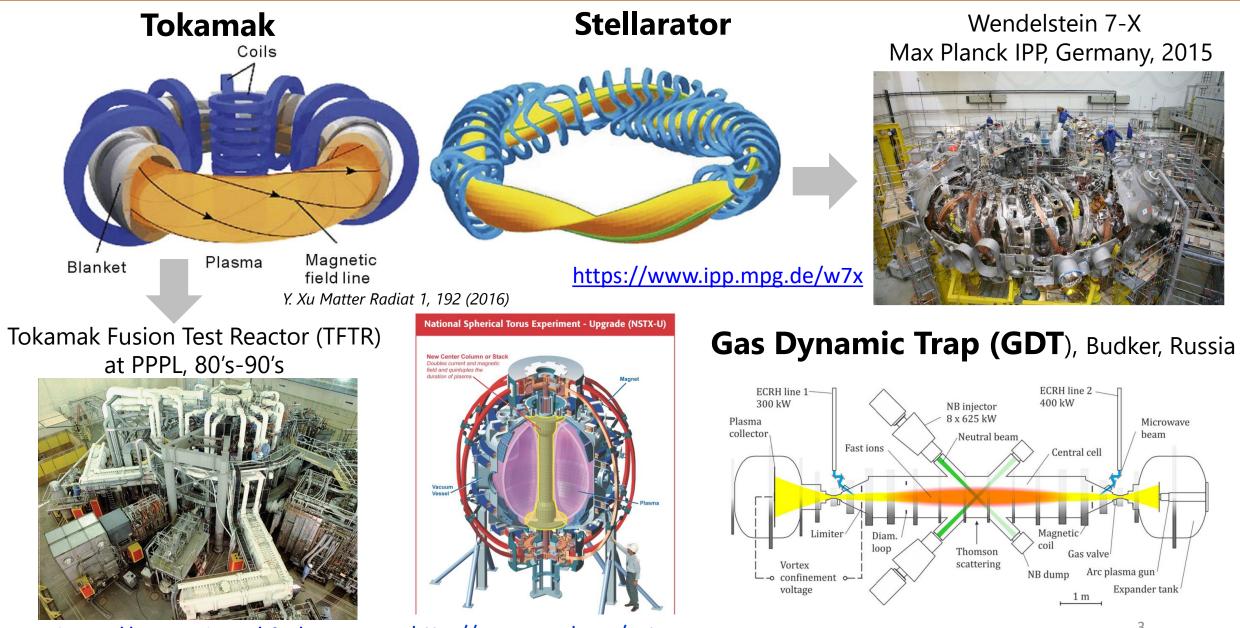
PPPL

GSS, Princeton, August 2020



- Magnetized plasma devices ( $r_L = \frac{mv}{qB} \ll L, \omega_c/v_t >>1$ )
- Magnetically-controlled LTPs for material processing at atomic scale
- Electromagnetic plasma propulsion: *Hall thrusters*

### Fusion plasma devices with magnetic confinement



https://w3.pppl.gov/tftr/

http://www.pppl.gov/nstx-u

A. A. Ivanov and V. V. Prikhodko, Plasma Phys. Control. Fusion **55** (2013)

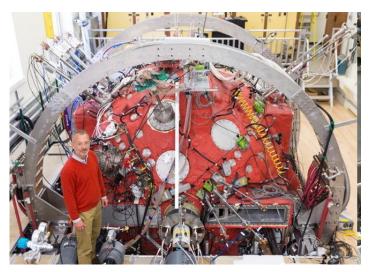
### Basic science experiments with magnetized plasmas

#### Large Area Plasma Device (LAPD)

UCLA, <a href="https://plasma.physics.ucla.edu/">https://plasma.physics.ucla.edu/</a>



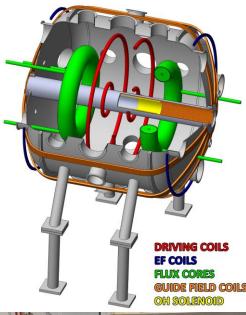
**Big Red Ball (BRB)** Wisconsin, <u>https://wippl.wisc.edu/big-red-ball-brb/</u>



Magnetized Dusty Plasma Experiment (MDPX) Auburn, <u>http://wp.auburn.edu/mprl/</u>

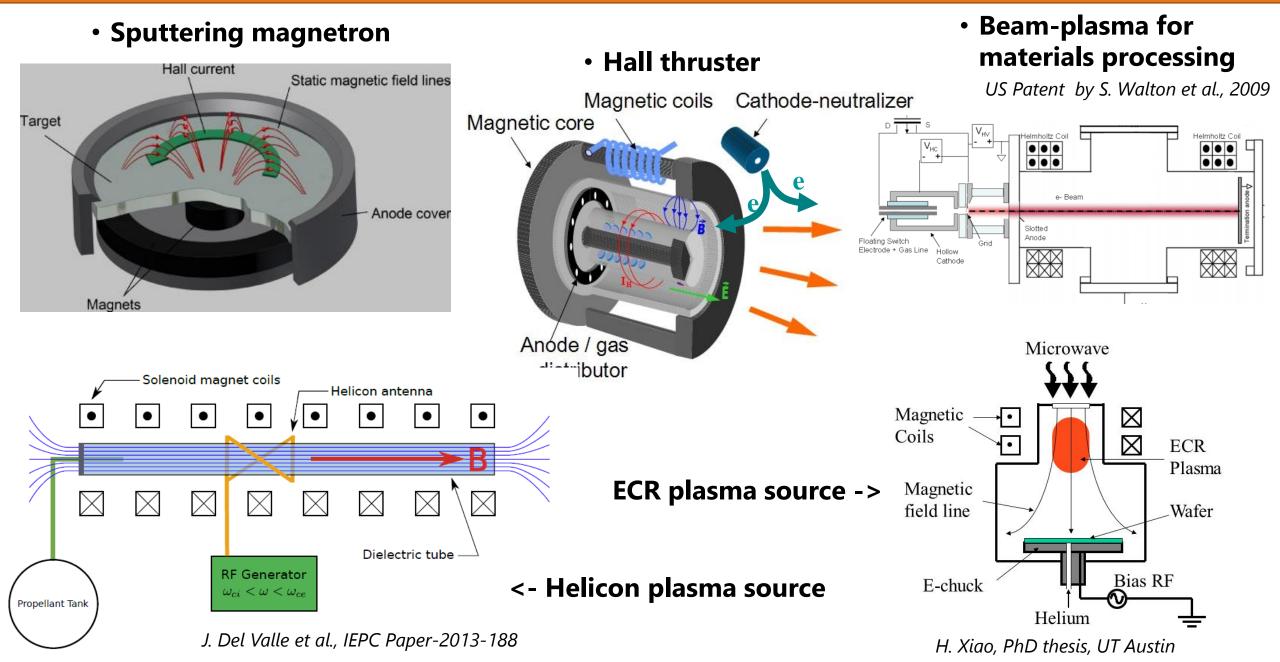


**FLARE** PPPL, <u>https://flare.pppl.gov/</u>





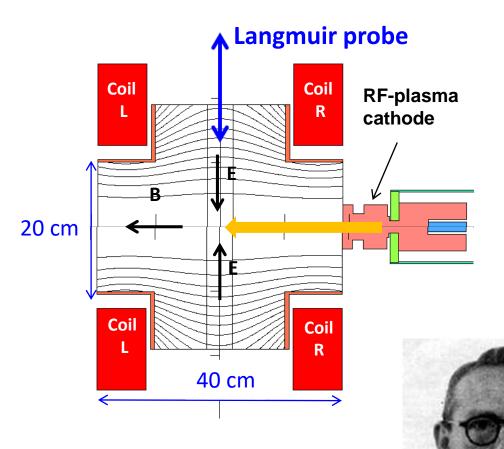
## Why to apply magnetic field in LTP?



## What is the use for the magnetic field in LTP?

• e-Beam plasma source





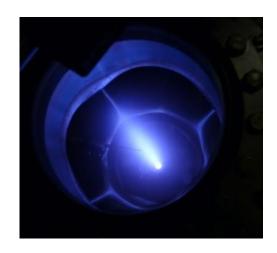
- Pressure: ~ mtorr
- Gas: 80% Ar+10% H<sub>2</sub>
- Magnetic field ~ 10-15 G
- Electron Beam
  - Energy  $< 10^2 \text{ eV}$
  - Current ~ 1 A

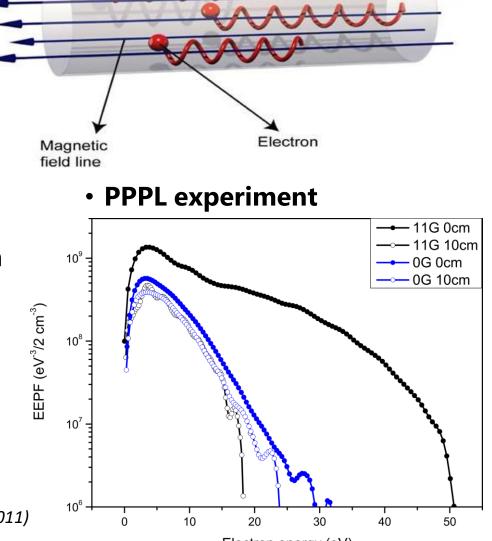
Frans Michel Penning Philips, 40's -?

# What to expect from the LTP with applied magnetic field?

With magnetic field

#### • e-Beam plasma source





#### Reduced transport across B-field

 $v_t \propto n_g = P_g / kT_g$  $(\Omega_{ec} / \nu_t)^2 >> 1$  $\mu_\perp \approx \frac{\mu}{(\Omega_{ec} / \nu_t)^2}$ 

 $D_{\perp} \approx \frac{D}{(\Omega_{ec}/\nu_t)^2}$ 

6-X higher plasma density and much more energetic electrons at the axis with B-field than without

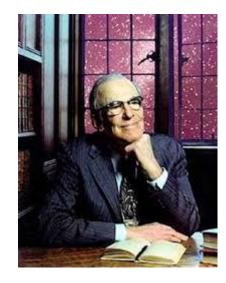
 Langmuir probe measured electron energy distribution function (EEDF)

$$N = \frac{2\sqrt{2m}}{|e|S_{\rm p}} \int_0^{-\infty} I_{\rm e}''(V)\sqrt{V/e} \,\mathrm{d}V$$

$$T_{\rm e} = \frac{4\sqrt{2m}}{3NS_{\rm p}} \int_0^{-\infty} I_{\rm e}''(V) (V/e)^{3/2} \,\mathrm{d}V.$$

V. Godyak, V. Demidov, J. Phys. D: Appl. Phys. 44 (2011)

### **Plasma propulsion research at PPPL**



#### Interplanetary Travel Between Satellite Orbits

By LYMAN SPITZER, JR.<sup>2</sup>

Princeton University Observatory, Princeton, N. J.

An analysis is given of the performance to be expected of a rocket powered by nuclear energy, and utilizing an electrically accelerated ion beam to achieve a gas ejection velocity of 100 km/see without the use of very high temperatures in the propellant gases. While such a rocket travel particularly feasible. It is well known that one of the chief limitations on a conventional rocket is the temperature which the rocket tubes can tolerate without melting or evaporating. A nuclear-powered

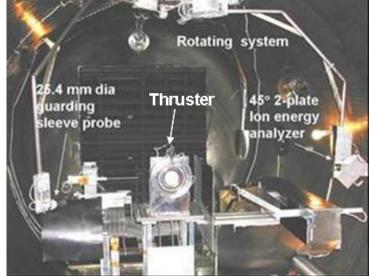
• Hall Thruster Experiment (HTX) in 1998 by N.J. Fisch and Y. Raitses

and graduate students L. Dorf, A. Smirnov, A. Litvak and D. Staack

• Goal: to develop scientific understanding of plasma thruster physics

2<sup>nd</sup> International Congress on Aeronautics, London, UK, 1952

1999

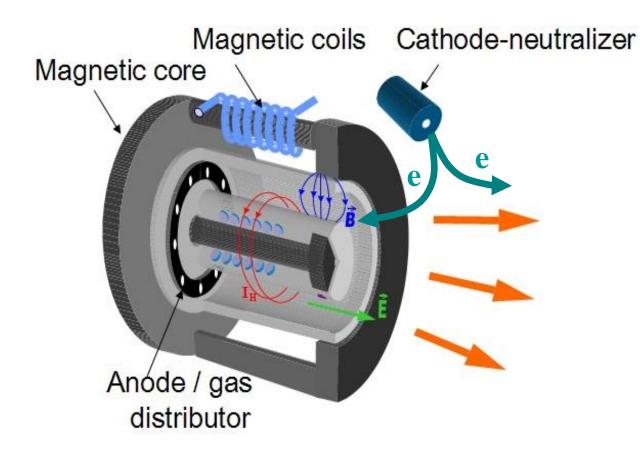


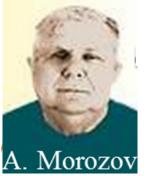


http://htx.pppl.gov



### Hall thruster invented in 60's, flown in 70's





Xenon, KryptonDimensions:V ~ 100-1000 VL < 10 cm</td>B ~  $10^2$  GaussD ~ 1-50 cmLow gas pressure < 1 mtorr</td>– ions move without collisions

- Applied DC (stationary) fields:  $\mathbf{E} imes \mathbf{B}$
- Quasineutral plasma:  $n_e \approx n_i$
- Electrons **ExB** drift in azimuthal direction
- Heavier ions almost unaffected by B-field

 $r_{Le} \ll L < r_{Li}$ 

• Equipotential magnetic field surfaces

 $\mathbf{E} = -\mathbf{V}_{\mathbf{e}} \times \mathbf{B}$ 

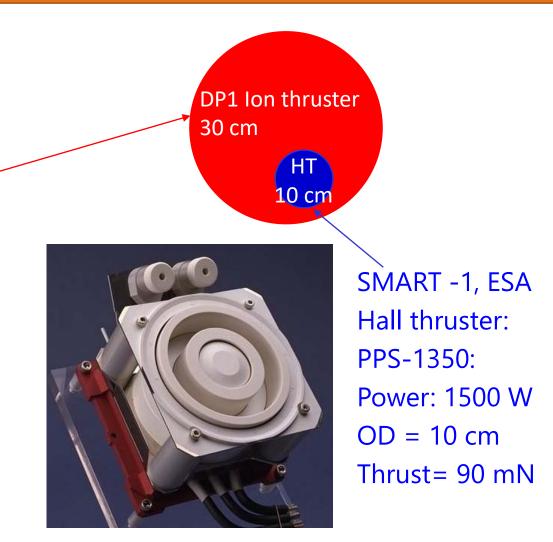
- Ions are accelerated by electric field
- Accelerated ion flux is neutralized by electrons
- Thrust force exerted on magnets

### **Comparing thrust densities for Hall and ion thrusters**

### Deep Space 1 NSTAR Ion thruster



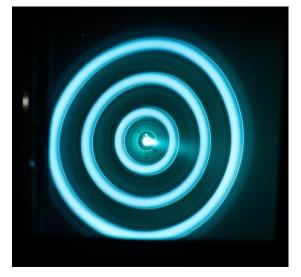
### At 1.5 kW, thrust $\approx$ 50 mN



Unlike ion thrusters, Hall thrusters are not space charged limited They have also a simpler design than ion thrusters Operational Satellites with Electric/ Plasma Propulsion (2008)



### World's largest and most powerful Hall thruster: X-3



Nested Hall thruster, X-3





Power	2-200 kW			
Efficiency	> 60%			
Thrust	0.2 N– 10 N			
lsp	1550 – 3500 s for <b>Xe</b>			
Diameter	0.80 m			
Thrust density	Up to ≈ <b>20 N/m</b> ²			
Mass	250 kg			

At higher power levels (over 600 kW, provided by a few X3s clustered together), the X3 has the potential to actually carry astronauts to Mars

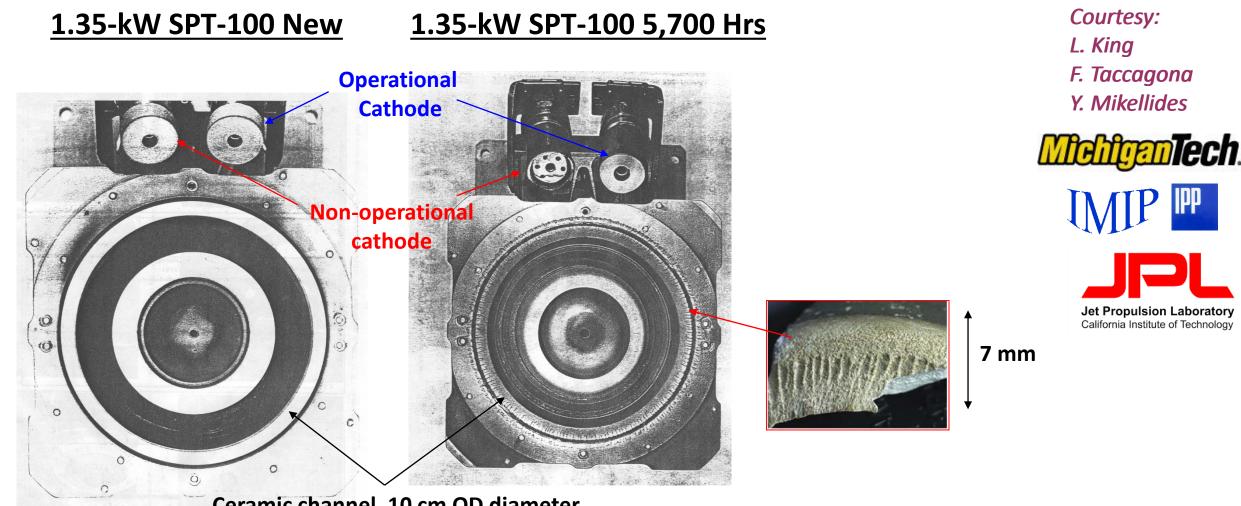
Courtesy Prof. Ben Jorns of Univ. Michigan Ann Arbor

## Key obstacles for the development of high thrust density and higher power Hall thrusters

- Ion-induced erosion of the thruster parts limiting the thruster lifetime
- Turbulent fluctuations and related anomalous transport phenomena, and plasma structures critically affecting the thruster performance and lifetime
- Enhanced plasma wall losses in scaled down miniaturized Hall thruster much less efficient low power thrusters than their larger counterparts
- Facility effects making difficult reliable testing of mid (10's kW) and high power (> 100 kW) thrusters



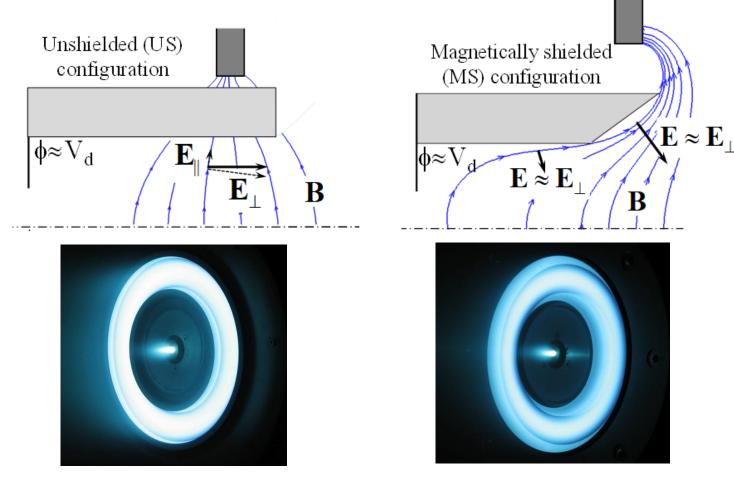
### **Ion-induced erosion of the thruster channel**



Ceramic channel, 10 cm OD diameter



### **Magnetically-shielded Hall thruster**

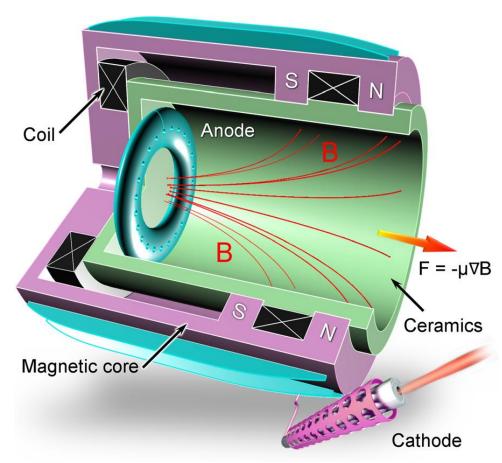


Courtesy: Vernon Chaplin Richard Hofer Yaingos Mikellides

Jet Propulsion Laboratory California Institute of Technology

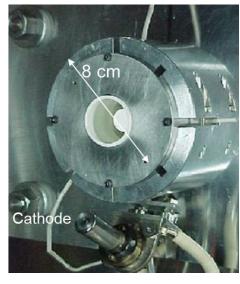
Magnetic shielding has been shown to dramatically reduce discharge channel wall erosion Central pole piece erosion is still a concern!

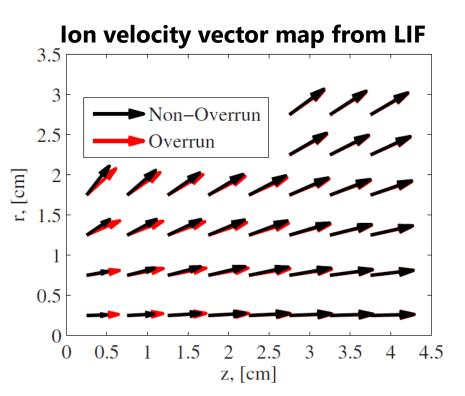
## **Cylindrical Hall thruster**





- No central channel wall
- Closed **E**×**B** drift (like in conventional Hall thruster)
- Electrons confined in a magneto-electrostatic trap
- Ion acceleration in a large volume-to-surface channel
- Plume focusing controlled by the cathode mode

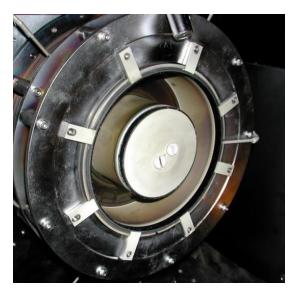




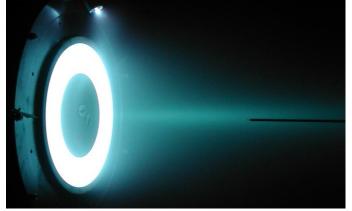
Y Raitses and N. J. Fisch, Phys. Plasmas, 8, 2579 (2001) A. Smirnov, Y. Raitses, and N.J. Fisch, Phys. Plasmas **14**, 057106 (2007) R. Spektor et al., Phys. Plasmas **17**, 093502 (2010)

### Plasma structures and anomalous cross-field current

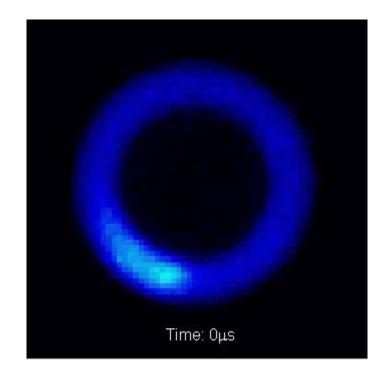
#### 12 cm diameter, 2kW Hall thruster



#### Xenon operation



**Plasma non-uniformity** Current conducting ExB rotating "spoke" Fast frame imaging 60 kfps



### **Thruster efficiency**

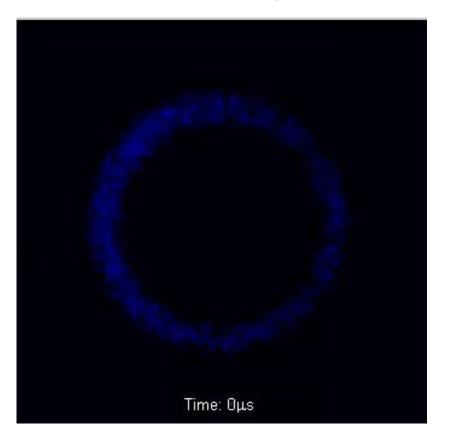
$$\eta \equiv \frac{TV_{jet}}{2P_e} \propto \frac{I_i}{I_i + I_{e\perp}}$$
  
Electron cross-field current

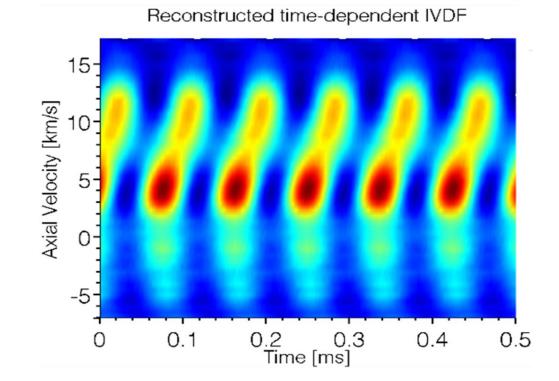
- Spoke frequency ~ 10 kHz
- 10's times slower than E/B
- Conduct > 50-70% of the discharge current

18

### Hall thruster DC inputs, but highly oscillatory operation

Breathing oscillations associated with ionization instability, 10-20 kHz





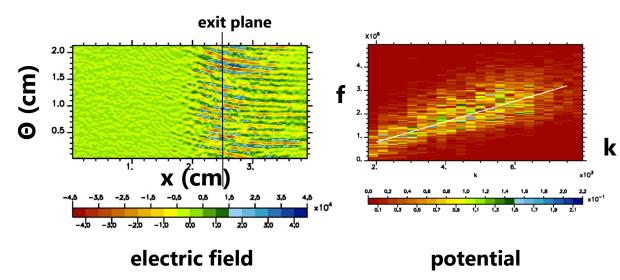
Powerful breathing oscillations of the discharge current due to ionization instability affects temporal evolution of ion velocity distribution



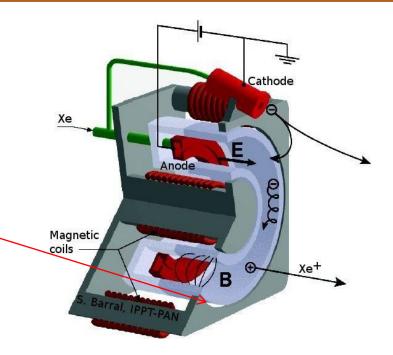
### **Small scale fluctuations**

#### **Outside the channel and in the near-field plume:** Small scale fluctuations ,~ 10 MHz, ion-sound and lower hybrid instabilities

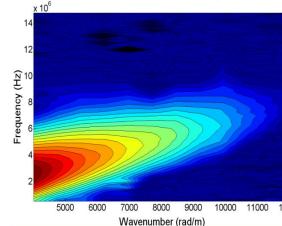
2-D self-consistent PIC simulations



Adam, Héron and Laval, Phys. Plasmas (2004) Tsikata, et al., Phys. Plasmas (2009)



• Experiment: Density fluctuations

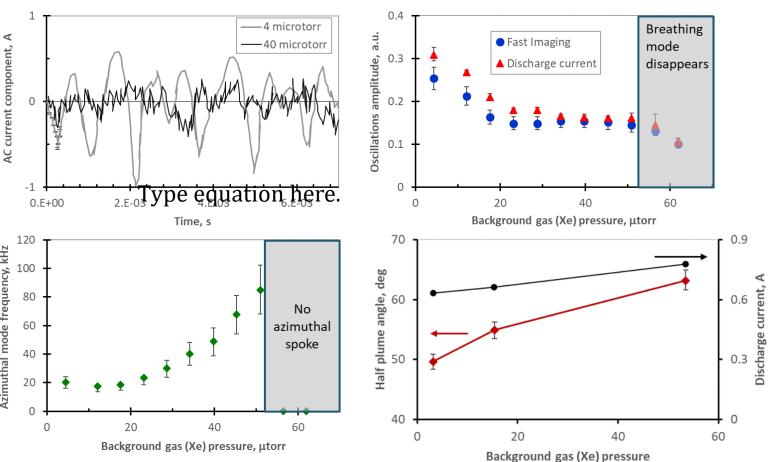


19

### **Facility effects**

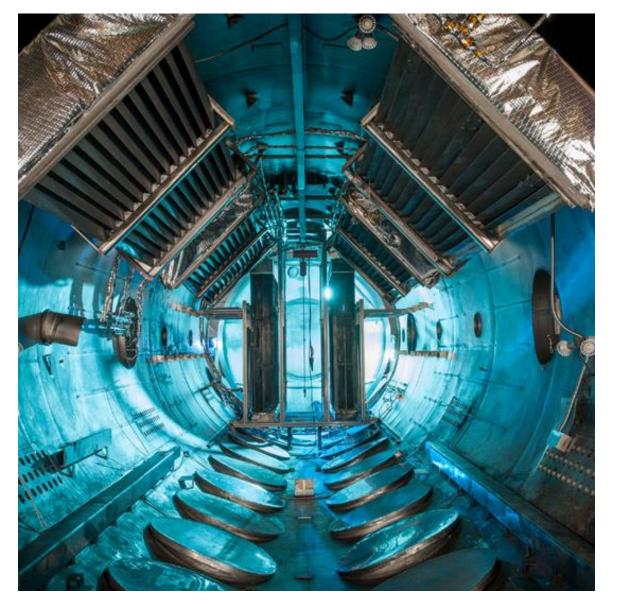
PPPL 28 m<sup>3</sup>, 66000 l/s for Xenon Base pressure: 10<sup>-8</sup> torr < 5 x 10<sup>-6</sup> torr for 200 W thruster





 $\frac{\lambda_{mfp}}{L} \gg 1$ 

### Testing capabilities for high power plasma propulsion



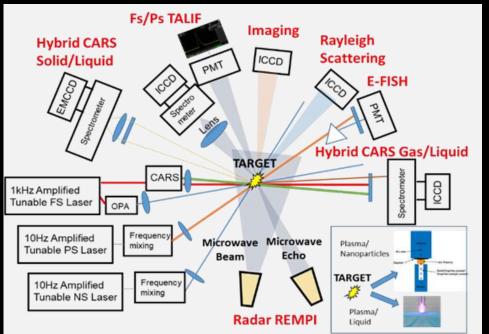
NASA's Glenn Research Center Vacuum Chamber 5 (4.6 m diam by 18.3 m long, pumping 700 kL/s) provides a testing environment for advanced Solar Electric Propulsion needed for future space missions

## **Critical challenges of modern Hall thruster technology**

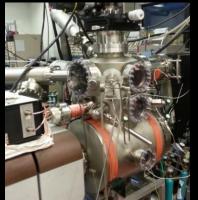
- How to make predictive Hall thruster designs for various power levels of interest (space mission relevant)?
  - Need experimentally validated predictive modeling capabilities (e.g. 3-D kinetic codes for plasma, atomistic simulations for materials)
- How to mitigate/suppress plasma instabilities?
  - Advance understanding of instabilities, develop active and passive engineering solutions (e.g. electrical circuitry, segmented electrodes)
- How to make accelerated thruster lifetime tests (e.g. 10's -100's hours instead of 1000's hours)?
  - Need experimental validated physics models and codes
- How to improve efficiency of miniaturized Hall thrusters
  - Need alternative operating regimes, new ExB configurations

# **PPPL-Princeton Collaborative LTP Research Facility (PCRF)**

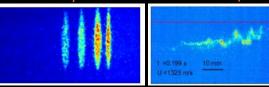
Advanced *fs-ps-ns-cw* diagnostics of plasma species, flow, nanoparticles



#### PFC properties



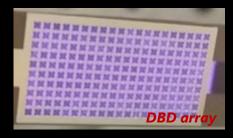
*fs*- Laser Electronic Excitation Tagging flow measurements
*Laminar flow* Turbulent flow



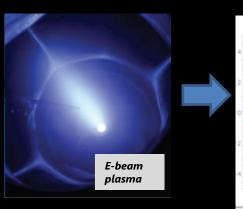
• ns-discharge in gas bubble

#### Plasma sources

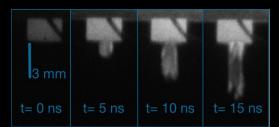
• Atmospheric pressure plasma: DBD, jets, arcs



• Low pressure LTPs: magnetized, e-beam, DC/RF microscale

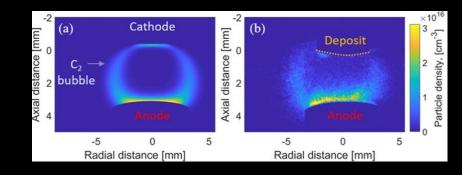


#### • Streamer in plasma jet

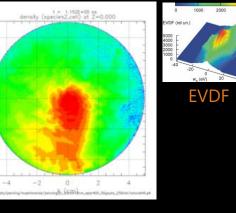


### Unique measurements/simulations

Plasmas with complex chemistry: arc discharge
2D CFD simulations (left) and LIF measurements (right)



#### • 1D, 2D, 3D Kinetic simulations



• Atomistic simulations



#### Nanoparticles synthesis in arc: *modeling*, OES and LII measurements

20 nm Nanoparticles growth region 15 nm Cathode deposit 10 nm 5 nm				
C <sub>2</sub> "bubble"				
1 2 3 4 s Anode	6	7	8	9
TiRe LII measurements : Signal integration region is highlighted	35±7 nm	25±5 nm	22±3 nm	

### **Princeton Collaborative Research Facility (PCRF)**

The PCRF mission is to provide the scientific community access to specialized, world-class diagnostics, instruments, computational tools, and related expertise

Upcoming Call for Collaborative Proposals

Opening call for proposals: November 2020

**Closing call for proposals: December 20, 2020** 

Review of proposals (tentative): December 21, 2020 to January 25, 2021

Deadline for decisions (tentative): by February 8, 2021





### Acknowledgement

- Jacob Simmonds, Fang Zhao, Ivan Romadanov, Andrei Smolyakov, Igor Kaganovich, Chris Tully, and Ahmed Diallo
- Research supported by US DOE and AFOSR