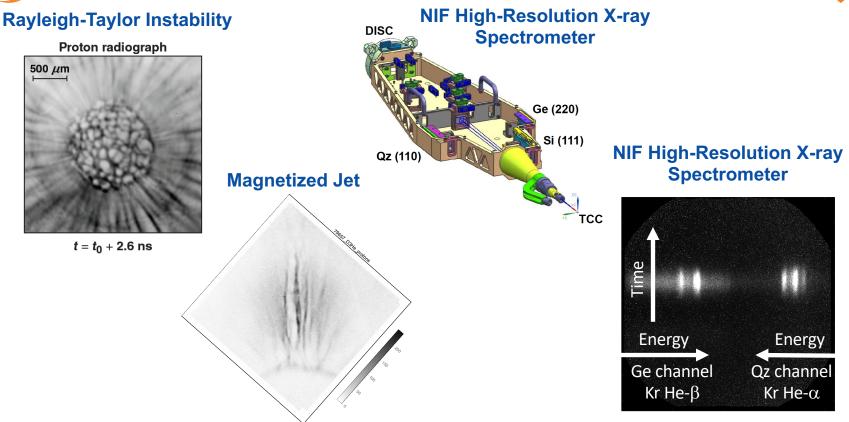
Laser-Driven High-Energy-Density Plasmas and Their Diagnostics







Lan Gao Princeton Plasma Physics Laboratory

PPPL Graduate Summer School August 12, 2019

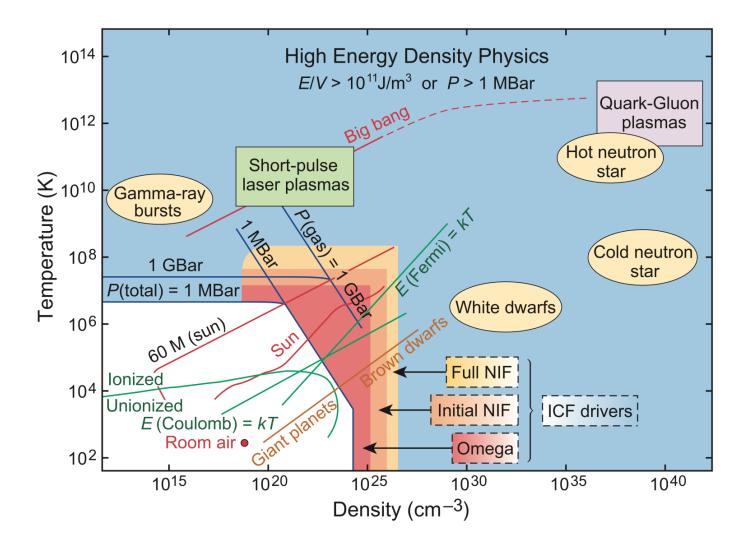
Outline





- Introduction
 - High energy density (HED) physics
 - Inertial confinement fusion
 - Laser-driven HED systems
 - Diagnostic requirement
- Basic diagnostic building blocks
 - Electromagnetic field
 - Particles
 - X rays
 - Pinhole imaging
 - Streak cameras
 - Framing cameras
 - Plasma conditions

High energy density (HED) physics concerns the study of matter at high densities and extreme temperatures*

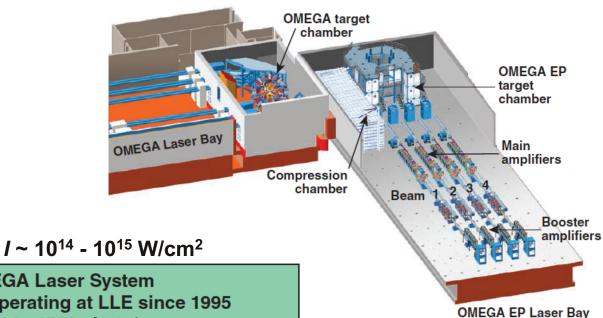


^{*} Frontiers in high energy density physics: The X-games of contemporary science. (The National Academies Press, Washington, DC, 2003).

Laboratory for laser energetics (LLE) at University of Rochester operates two of the world's large lasers for HED physics research







 $I \sim 10^{18} \text{ W/cm}^2$

OMEGA Laser System

- Operating at LLE since 1995
- Up to 1500 shots/year
- **Fully instrumented**
- 60 beams
- >30-kJ UV on target
- 1% to 2% irradiation nonuniformity
- Flexible pulse shaping
- Short shot cycle (1 h)

More than half of OMEGA's shots are for external users.

OMEGA EP Laser System

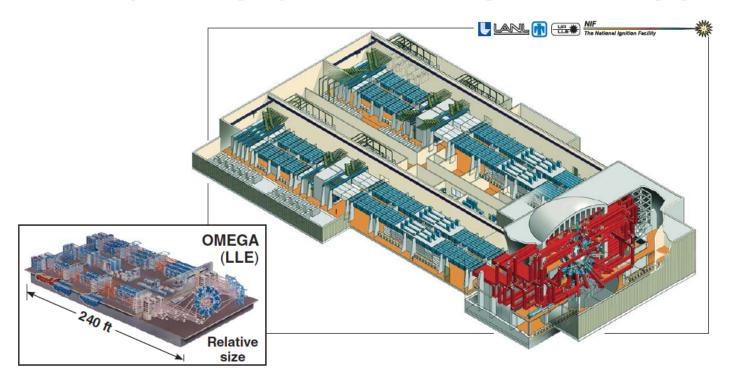
- Completed 25 April 2008
- Four NIF-like beamlines; 6.5-kJ UV (10 ns)
- Two beams can be high-energy petawatt
 - 2.6-kJ IR in 10 ps
 - can propagate to the OMEGA or **OMEGA EP target chamber**



The National Ignition Facility (NIF) at LLNL aims at demonstrating fusion ignition



- The NIF is a 1.8-MJ laser system (60× OMEGA's energy);
 NIF is a \$3.5 billion facility completed in 2009
- The NIF is performing experiments with the goal of achieving ignition



The achievement of ignition—a national "grand challenge"—on the NIF will change the fusion landscape.

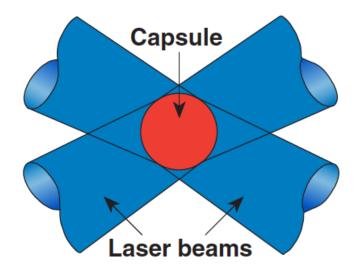
* How NIF works, https://www.youtube.com/watch?v=yixhyPN0r3g

Both direct and indirect (x-ray) drive are being used to implode the inertial confinement fusion capsules

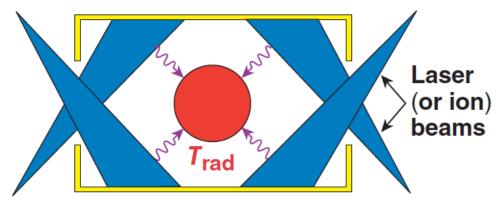




Direct-drive target



X-ray-drive target



Hohlraum using a cylindrical high-Z case $T_{\rm rad}$ is the x-ray temperature

Key physics issues

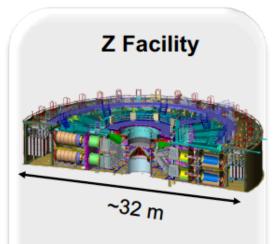
- Energy coupling
- Drive uniformity
- Hydrodynamic instabilities
- Compressibility



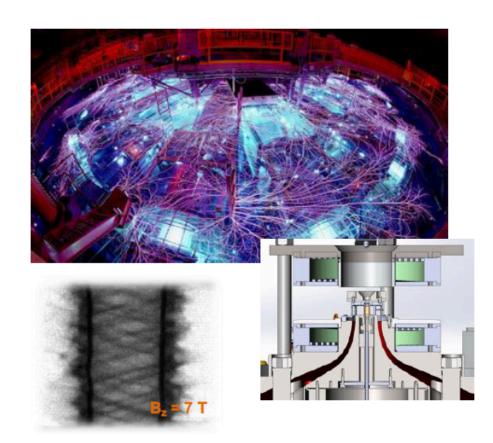
Magnetic drive ICF is being pursued at the Z pulsed power facility at Sandia National Laboratories, NM





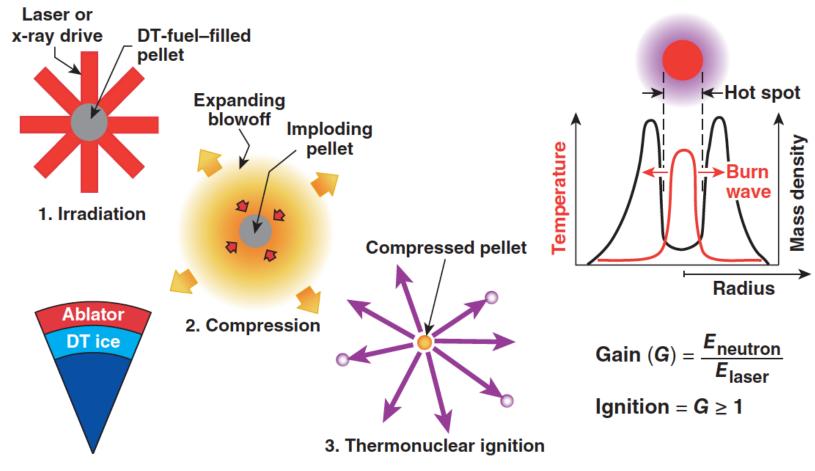


- 80 TW peak electrical power
- Up to ~1 MJ of electrical energy
- Optimized for magnetic drive



Ablation is used to generate the extreme pressures required to compress a fusion capsule to ignition



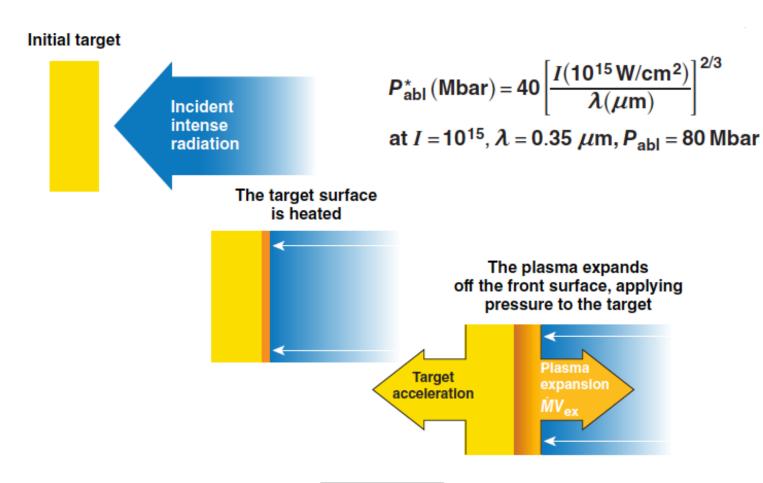


"Hot-spot" ignition requires the core temperature to be at least 5 keV and the core fuel areal density to exceed ~300 mg/cm².

Intense lasers create HED conditions in the laboratory through ablation





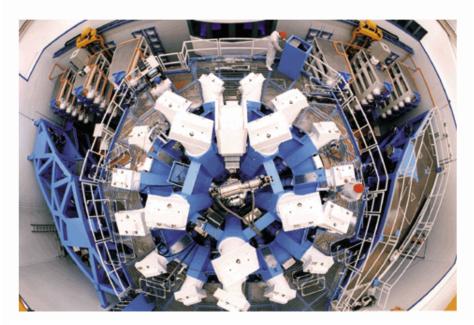


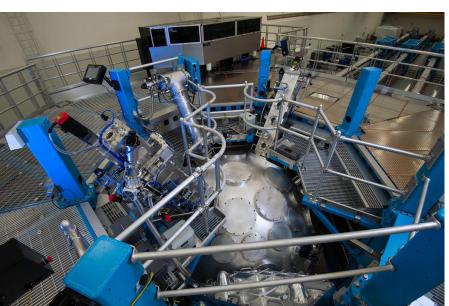
^{*}J. D. Lindl, Inertial Confinement Fusion: The Quest for Ignition and Energy Gain Using Indirect Drive (Springer-Verlag, New York, 1998).

If you look at the OMEGA target chamber







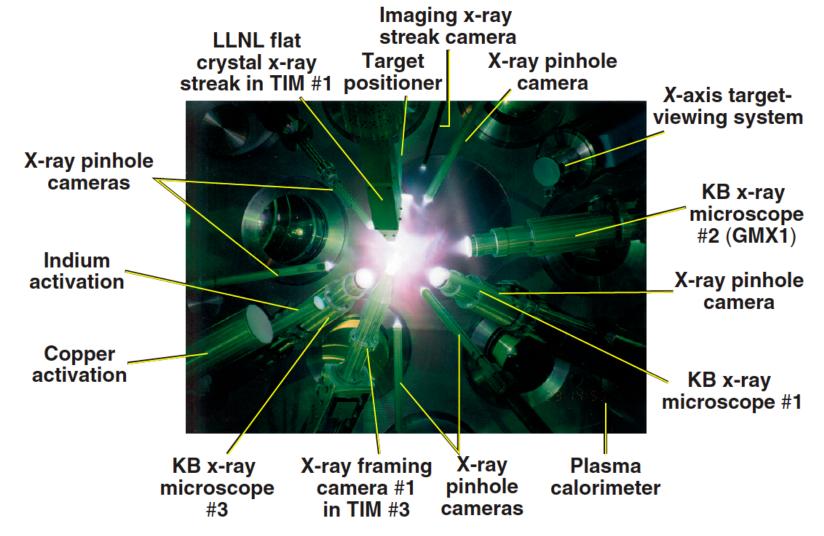


OMEGA EP

If you look at the OMEGA target chamber



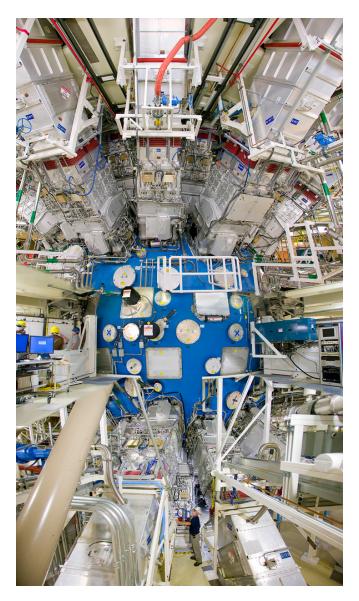


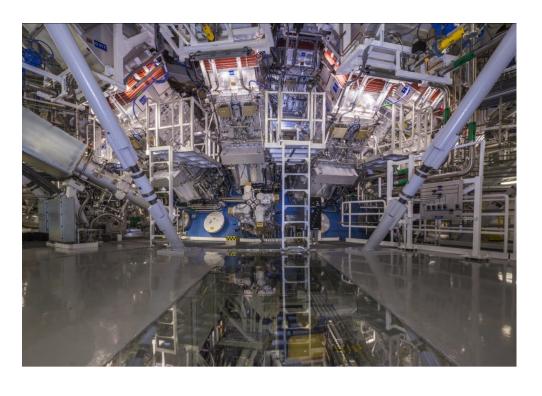


The NIF target chamber





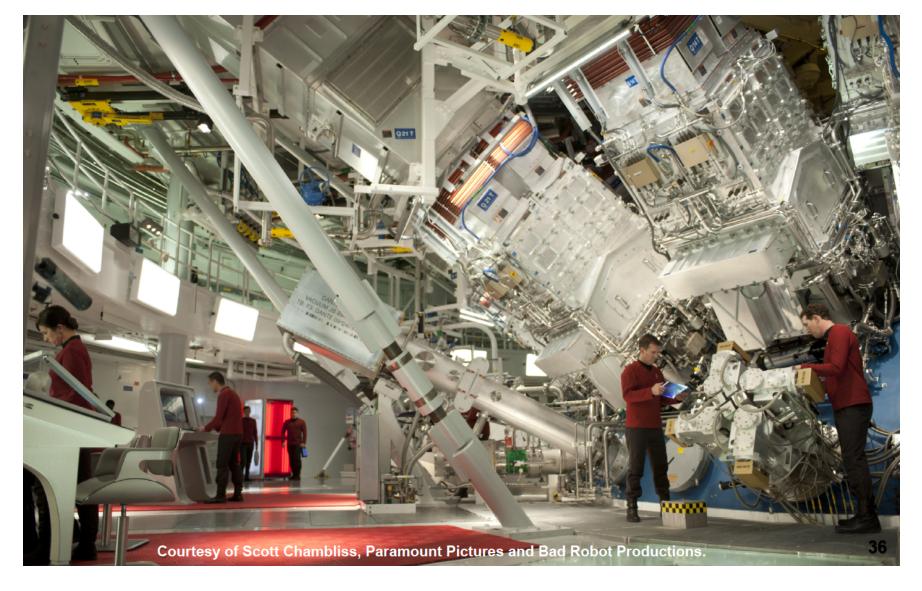




The NIF target chamber







High-energy-density-physics (HEDP) systems are diagnosed by optical, x-ray, particle and nuclear means

- HEDP systems generate some or all of
 - Visible light
 - UV and x-ray photons
 - Charged particles
 - Neutrons
 - Strong fields
- A comprehensive diagnostic suite makes it possible to learn a great deal about the systems: field strength and their impact, plasma parameters (n_e, n_i, T_e, v), particles, instabilities, yield, etc...
- Diagnosing HED systems require very high temporal (sub-ns, ps)
 and spatial (~10 μm) resolution

Diagnostic performance is determined by the resolution and signal-to-noise levels





- Spatial, temporal, or energy resolution determines the diagnostic properties
- The resolution depends on the design and on the signal-tonoise (background) ratio
- The signal level depends on
 - Source brightness
 - Solid angle of the detector $\Delta\Omega$ = $A_{det}/4\pi D^2$, where A is the effective diagnostic area, and D is the distance from the source to the diagnostic
- The noise (background) level is determined by design and intrinsic noise level (e.g., photon statistics)
- For example, when low number of particles (N) are detected, the uncertainty scales as sqrt(N)/N

Outline





- Basic diagnostic building blocks
 - Electromagnetic field
 - Proton radiography
 - Particles
 - RCF stack / Proton activation pack / Electron spectrometer / Thomson parabola
 - X rays
 - Pinhole camera (time integrated) / 1-D streak camera / 2-D framing camera
 - Plasma conditions
 - High-resolution x-ray spectroscopy / Thomson scattering / Neutrons / x-ray radiography

Outline





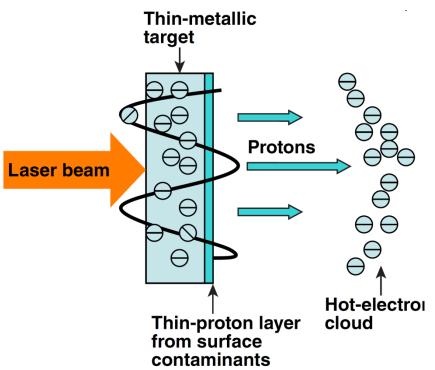
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Target Normal Sheath Acceleration (TNSA)* generates MeV proton beams in intense (>10¹⁸ W/cm²) laser-solid interactions





- Hot electrons escape from the rear side of the target
- An electrostatic field is built up, with a field gradient of the order of MeV/ μ m
- Protons are accelerated to tens of MeV

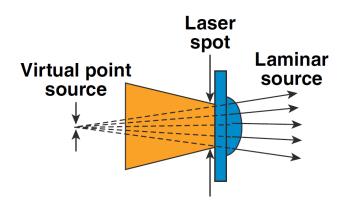


Laser-driven protons are ultra bright, extremely collimated, and have high peak energy (58 MeV) and short burst duration (picosecond scale).

^{*}S. C. Wilks et al., Phys. Plasma 8, 542 (2001)

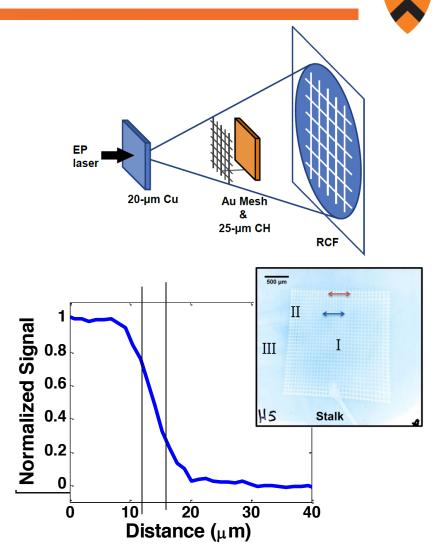
The virtual proton source is much smaller than the laser spot*





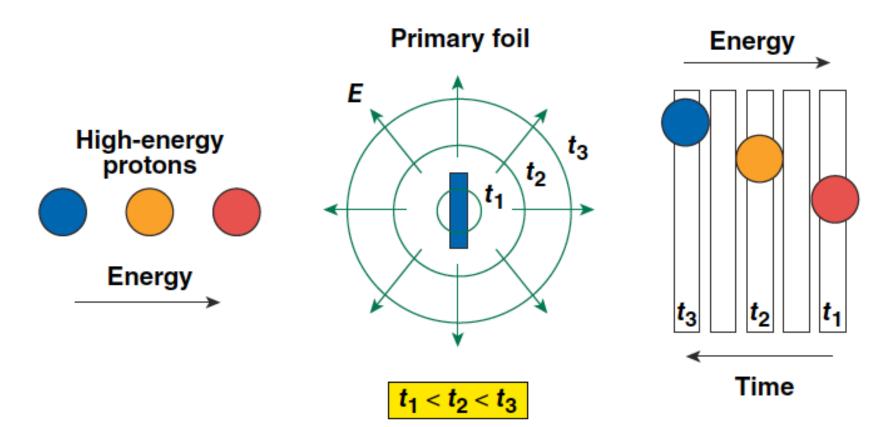
Virtual source size: $4\pm2~\mu m$

Very high spatial resolution!



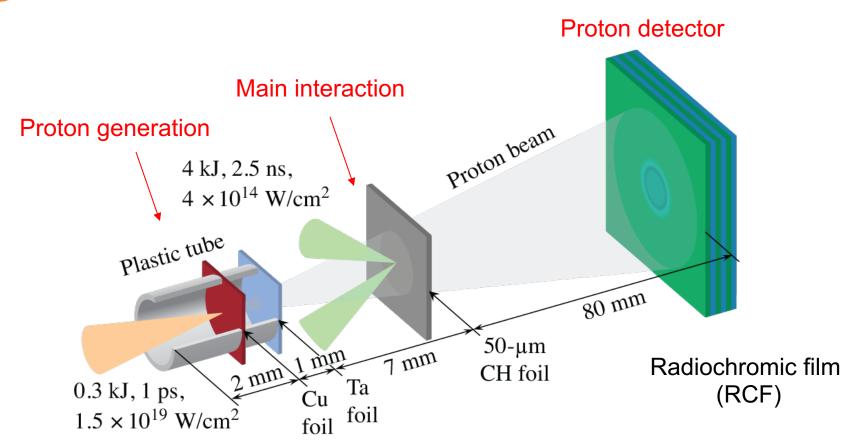
~ 8 µm for 15 MeV protons

Time-of-flight dispersion and a filtered stack detector produces a multiframe imaging capability



Ultrafast laser-driven proton radiography is developed on OMEGA EP

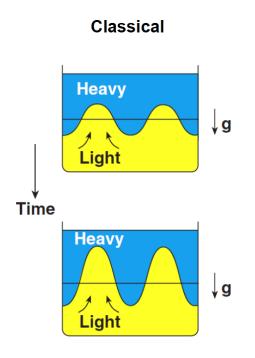




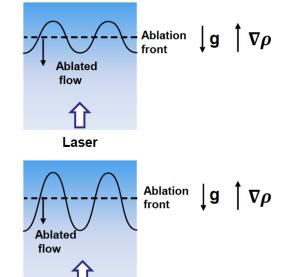
The laser-driven target is subjected to the Rayleigh-Taylor (RT) instability

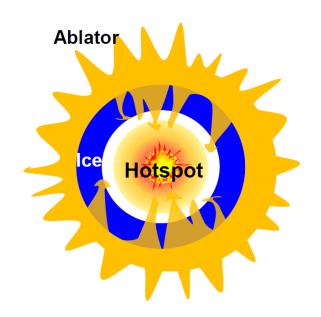












The RT instability has linear and nonlinear stages*





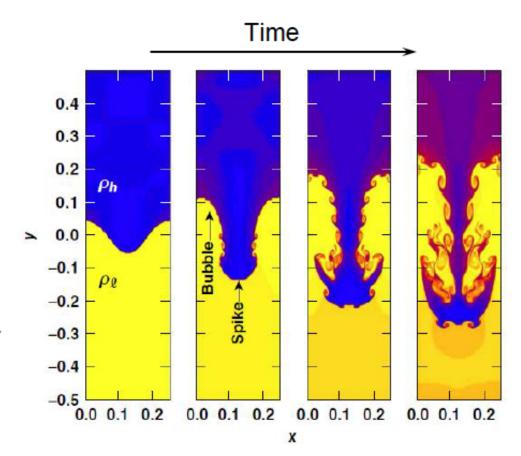
Linear regime (classical)**:

$$\eta = \eta_0 e^{\gamma t}$$
 $\gamma = \sqrt{AKg}, \quad A = \frac{\rho_h - \rho_l}{\rho_h + \rho_l}$

Linear regime (ablative)**:

$$\gamma = \alpha \sqrt{\frac{\kappa_g}{1 + \epsilon \kappa L}} - \beta \kappa V_a$$

• Nonlinear regime***: $\eta \geq 0.1\lambda$ Slower growth Bubbles and spikes Bubble competition and merger



^{*}Shengtai Li and Hui Li. "Parallel AMR Code for Compressible MHD or HD Equations". Los Alamos National Laboratory (2006).

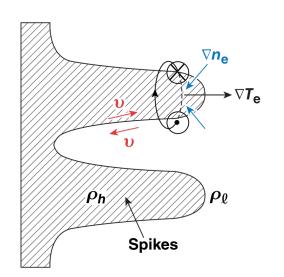
^{**}J. D. Kilkenny et al., Phys. Plasmas 1, 1379 (1994).

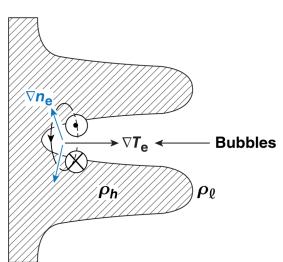
^{***} R. Betti and J. Sanz, Phy. Rev. Lett. 97, 205002 (2006).



Magnetic fields are generated by the Biermann battery mechanism







Biermann battery:

$$\mathbf{E} = -\frac{\nabla P_{e}}{\mathbf{e}n_{e}}$$

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E} \sim -\frac{\nabla n_{\mathrm{e}} \times \nabla T_{\mathrm{e}}}{\mathbf{e} n_{\mathrm{e}}}$$

Azimuthal magnetic fields are generated by $\nabla n_e \times \nabla T_{e.}$

^{*}K. Mima et al., Phys. Rev. Lett., 41, 1715 (1978);

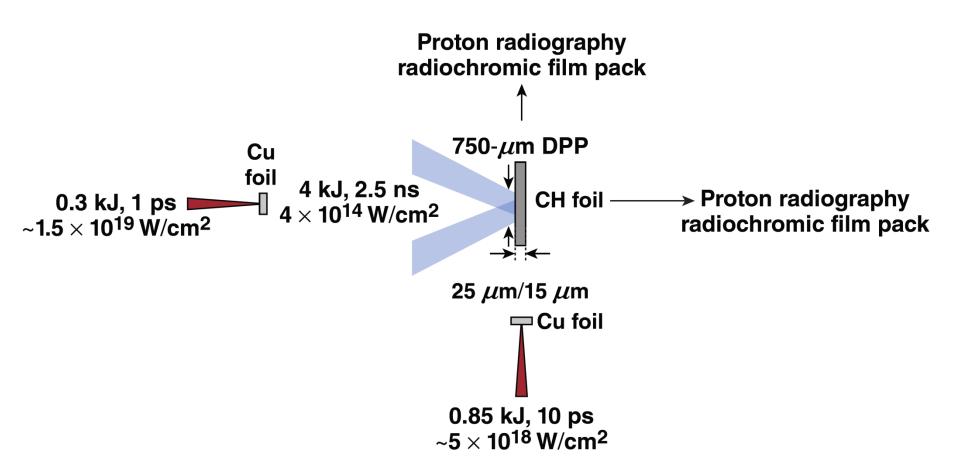
R. G. Evans, Plasma Phys. Control. Fusion., <u>28</u>, 1021 (1986);

B. Srinivasan et al., Phys. Rev. Lett., 108, 165002 (2012).

M. Manuel et al., Phys. Rev. Lett., 108, 255006 (2012).

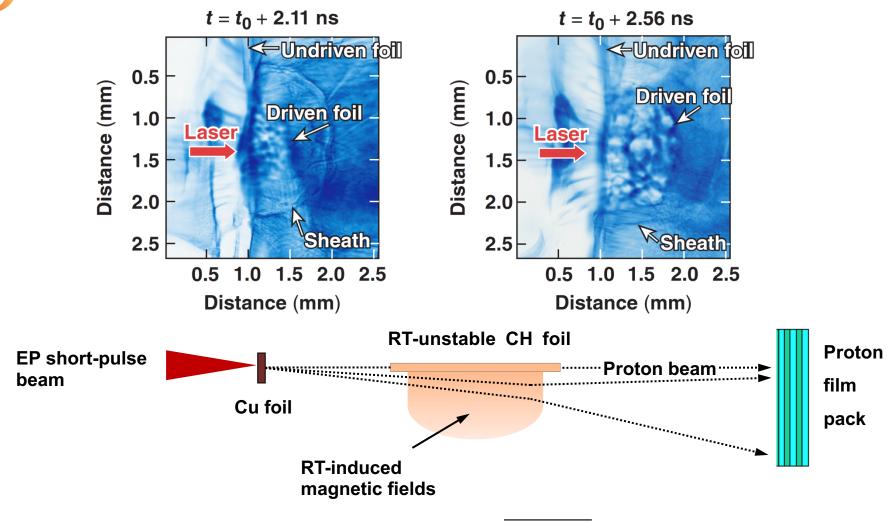
Magnetic-field generation has been studied in side-on and face-on geometries using the acceleration of planar plastic targets





Proton radiography of 15- μ m-thick foils reveals magnetic field generation and its evolution*



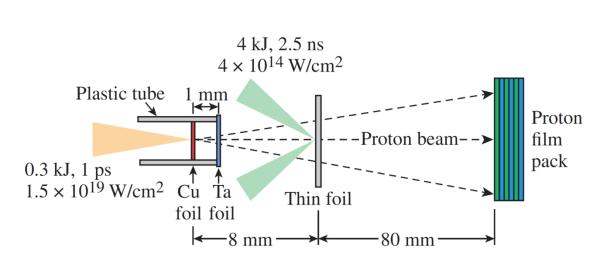


^{*}L. Gao et al., Phys. Rev. Lett. <u>109</u>, 115001 (2012).

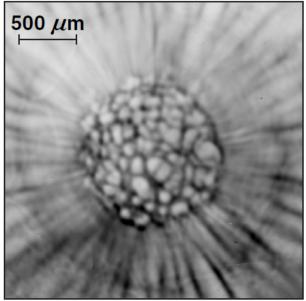
Face-on geometry

Face-on probing reveals magnetic field generation by the RT instability





Proton radiograph

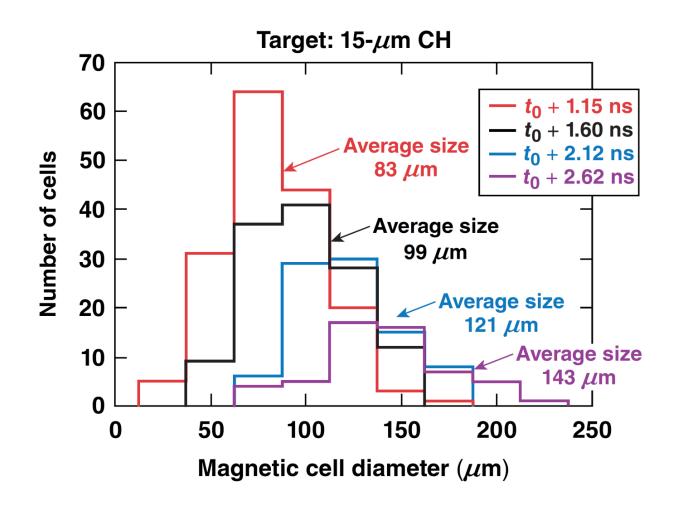


 $t = t_0 + 2.6 \text{ ns}$

The number of magnetic cells decreases and the magnetic cell diameter increases with time



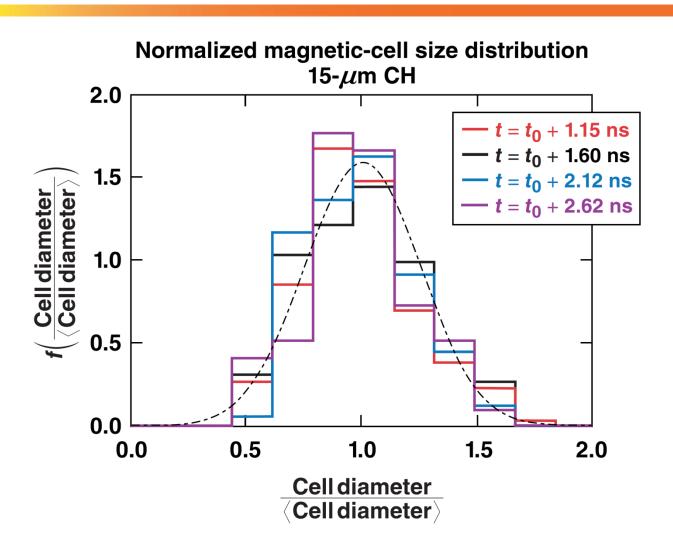




Face-on geometry

The normalized magnetic-field spatial distribution evolves self-similarly

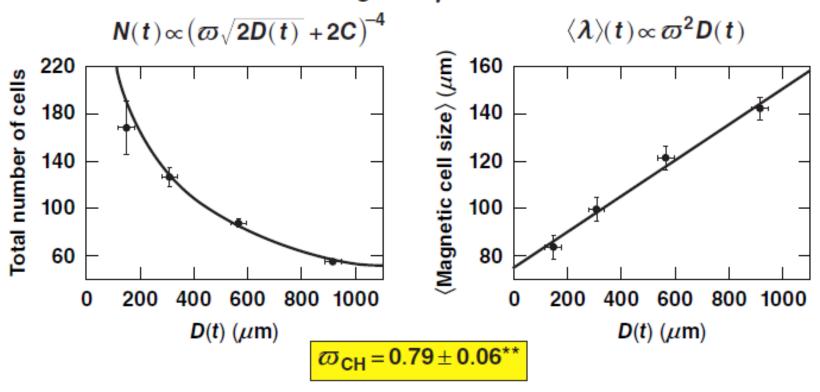




The evolution of the magnetic-field spatial distribution is consistent with an RT bubble competition and merger model*







^{*}O. Sadot et al., Phys. Rev. Lett. 95, 265001 (2005);

D. Oron et al., Phys. Plasmas 8, 2883 (2001);

U. Alon et al., Phys. Rev. Lett. 72, 2867 (1994).

^{**}L. Gao et al., Phys. Rev. Lett. 110, 185003 (2013).

The origin and amplification of the magnetic field in the universe is a central astrophysical problem





Sources of magnetic fields

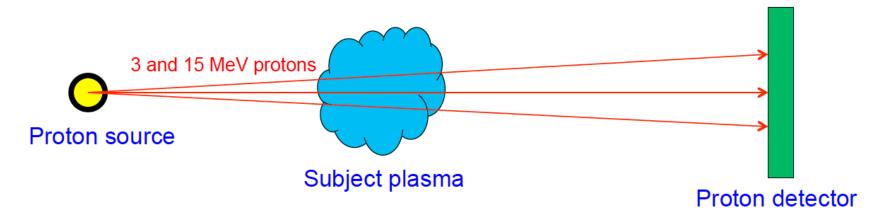
 Amplification by the dynamo process

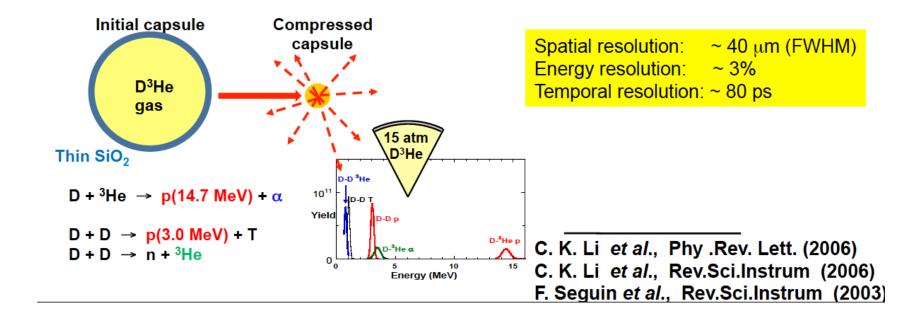
 Flow-dominated systems are common in astrophysics

Particle acceleration, nonthermal emission

A D³He mono-energetic proton radiography platform has been developed by MIT for HEDP experiments



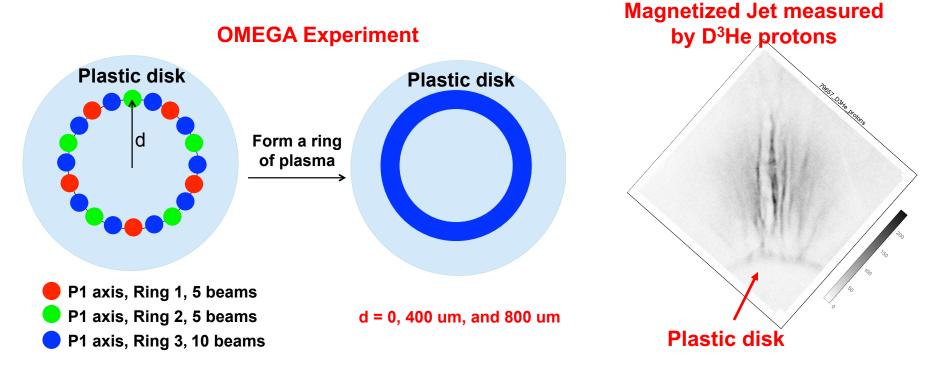




A strong, fast propagating, magnetized jet has been created at the OMEGA Laser Facility





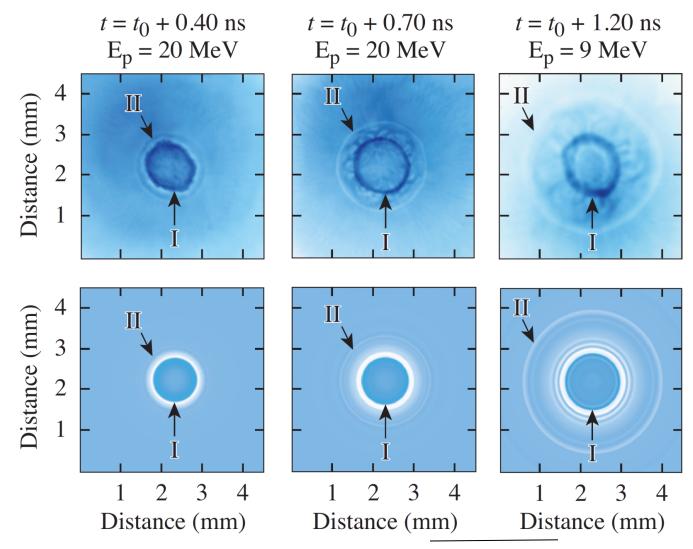


- Collisionless shocks are believed to sites for cosmic ray acceleration
- A magnetized, supersonic jet has been successfully demonstrated*
- The next goal is to collide the jets for collionless shock

^{*}L. Gao et al., The Astrophysical Journal Letters, 873:L11, 2019

Precision mapping of the laser-driven magnetic fields.



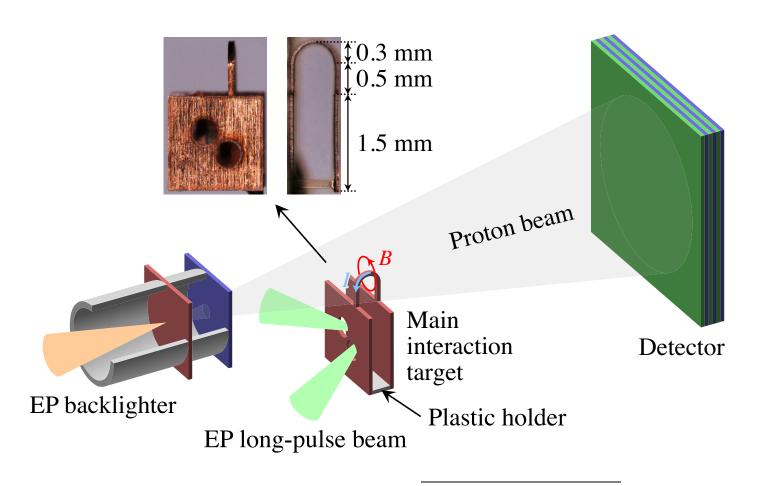


^{*}L. Gao et al., Phys. Rev. Lett. <u>114</u>, 215003 (2015).

Ultrafast proton radiography directly measured 100s of Tesla magnetic fields generated by a laser-driven capacitor coil target







^{*}L. Gao et al., Phys. Plasma 23, 043106 (2016)

Outline





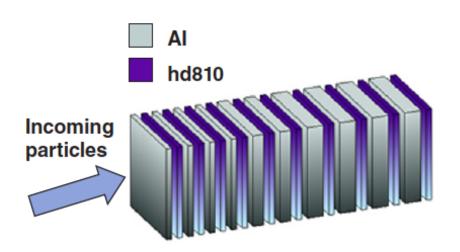
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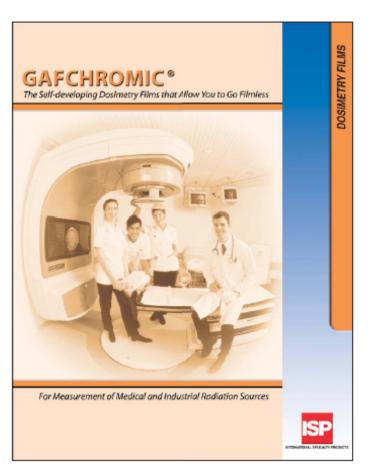


Radiochromic film allows simple detection of high-energy photons and particles



- Radiochromic film is self-developing and insensitive to visible light
- Multiple layers of film and filters allow energy discrimination



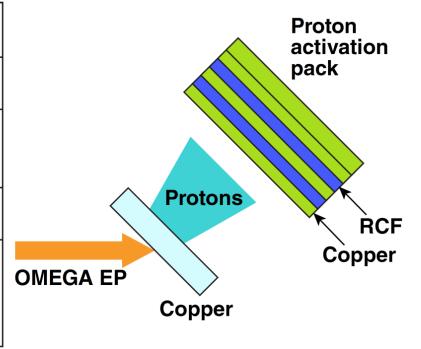




Experiments were performed on OMEGA EP to characterize energetic protons



Focal spot (R ₈₀)	20 ~ 25 <i>μ</i> m
Laser energy	40 J ~ 1500 J
Intensity (average within R ₈₀)	$0.25 \sim 8 \times 10^{18}$ W/cm ²
Intensity contrast	~108
Targets	$500~\mu\mathrm{m}^2 imes 20~\mu\mathrm{m}$ Cu/Cu+Al/Cu+CH $500~\mu\mathrm{m}^2 imes 50~\mu\mathrm{m}$ Cu

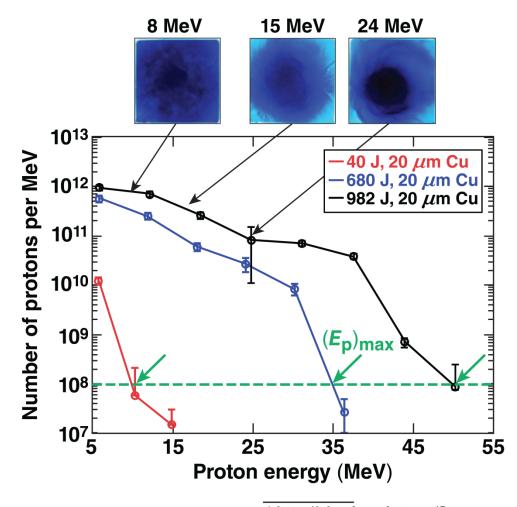


(1)

Nuclear activation of copper stacks determined the energy spectrum of the forward-accelerated protons



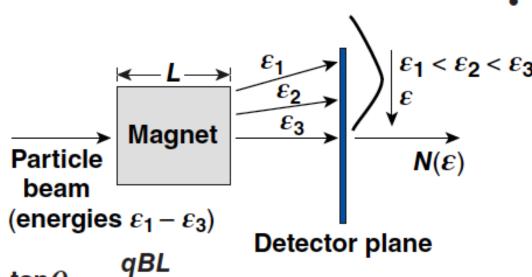
- Radiochromic film shows proton beam profile
- ⁶³Cu (p, n) ⁶³Zn
 ⁶⁵Cu (p, 3n) ⁶³Zn
- Coincidence counter absolutely calibrated using known source Na₂₂
- Response functions using stopping power* and cross-section data**
- An iterative method to recover the energy spectrum



^{*} http://physics.nist.gov/Star.

^{**} http://www.nndc.bnl.gov/exfor/.

The energy spectrum of high-energy charged particles can be measured with a magnetic spectrometer

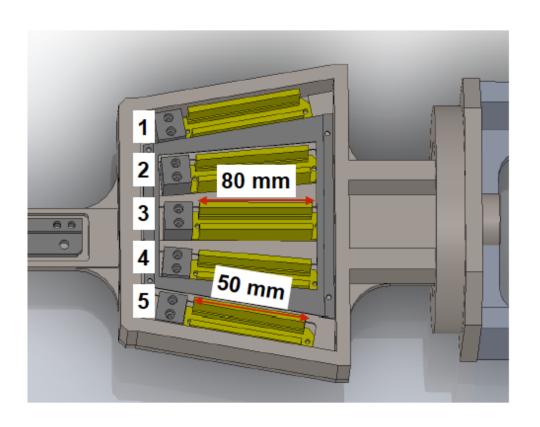


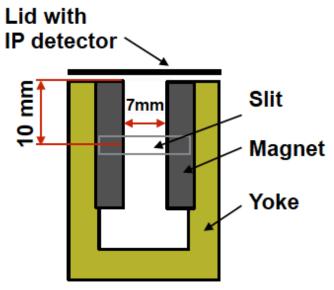
- Some issues
 - must know the detector response as a function of $oldsymbol{arepsilon}$
 - $-\frac{\Delta \varepsilon}{\varepsilon}$ is smaller at higher
 - the design can be optimized
 - magnet geometry
 - detector-plane orientation
- A major limitation is that the spectrometer cannot resolve among particles with the same sign of charge but different q, m, ε degenerate
 - A magnetic spectrometer works well for
 - electrons, positrons
 - ions when only a single species is present—typically protons

1

The energy spectrum of high-energy charged particles can be measured with a magnetic spectrometer

OU-ESM

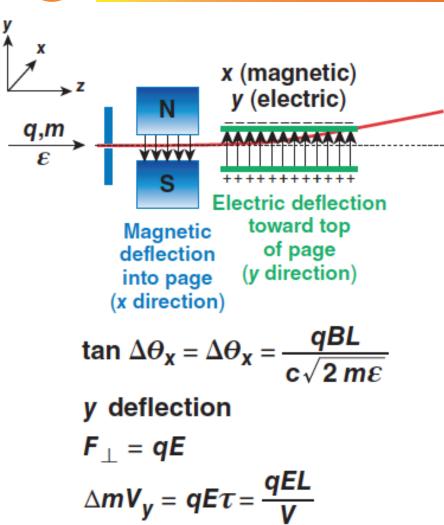




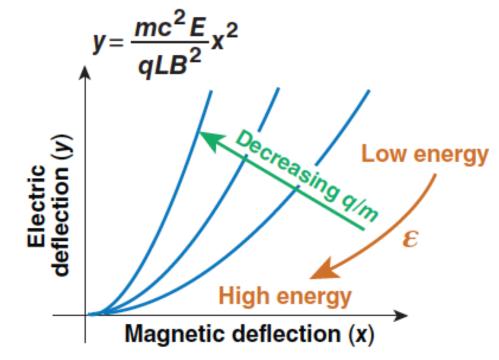
Ch1 -10° 50 mm length Ch2 -5° 80 mm length Ch3 -0° 80 mm length Ch4 +5° 80 mm length Ch5 +10° 50 mm length

A Thomson parabola uses parallel electric and magnetic fields to deflect particles onto parabolic curves that resolve q/m





- Deflection caused by magnetic field ~q/p
 - Deflection caused by electric field ~q/KE
- Ion traces form parabolic curves on detector plane



Outline





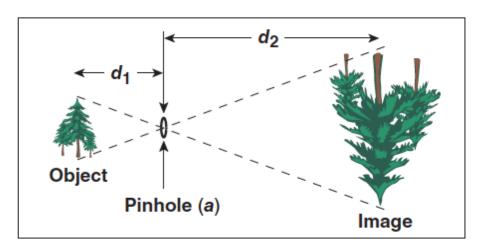
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The simplest imaging device is a pinhole camera



Kodak Brownie camera





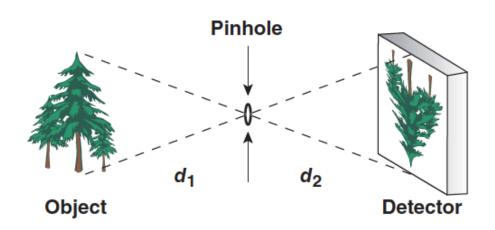
- Magnification = $\frac{d_2}{d_1}$
- Infinite depth of field (variable magnification)
- Pinhole diameter determines
 - resolution ~a
 - light collection: $\Delta\Omega = \frac{\pi}{4} \frac{a^2}{d_1^2}$

Imaging optics (e.g., lenses) can be used for higher resolutions with larger solid angles.

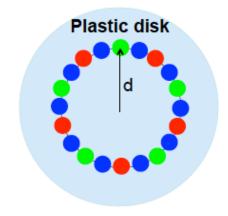
2-D time-integrated images can be recorded on film or electric detectors



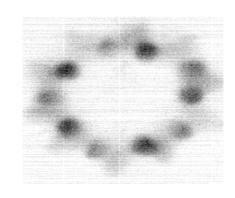




Magnetized Jet Exp

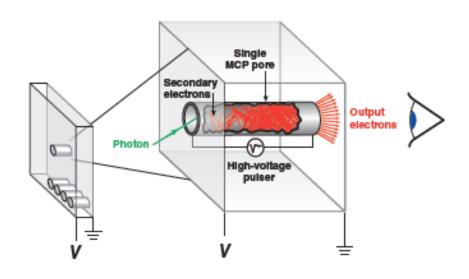


- A 2-D detector is required for a pinhole camera
 - film-requires processing
 - electronic
 - semiconductor arrays—signal proportional to incident flux per pixel (CCD or CID)
 - array of ionization detectors
 - single-photon counters—limited dynamic range



A framing camera provides a series of time-gated 2-D images, similar to a movie camera

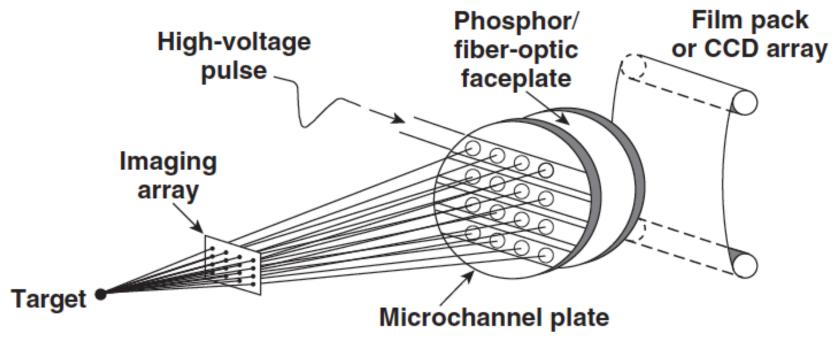
- The building block of a framing camera is a gated microchannelplate detector (MCP).
- An MCP is a plate covered with small holes.



Multiple electrons are produced each time an electron or photon hits the wall.

Two-dimensional time-resolved images are recorded using x-ray framing camera



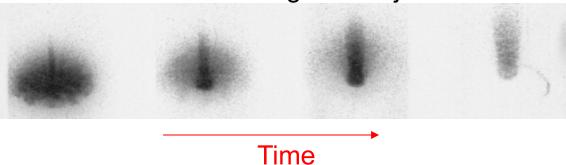


- Temporal resolution = 35 to 40 ps
- Imaging array: Pinholes: 10- to 12- μ m resolution, 1 to 4 keV
- Space-resolved x-ray spectra can be obtained by using Bragg crystals and imaging slits

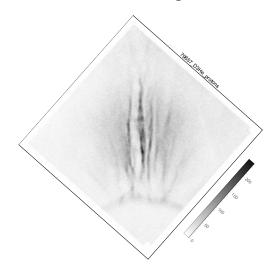
Two-dimensional time-resolved images are recorded using x-ray framing camera



X-ray framing camera: self-emission of the magnetized jet



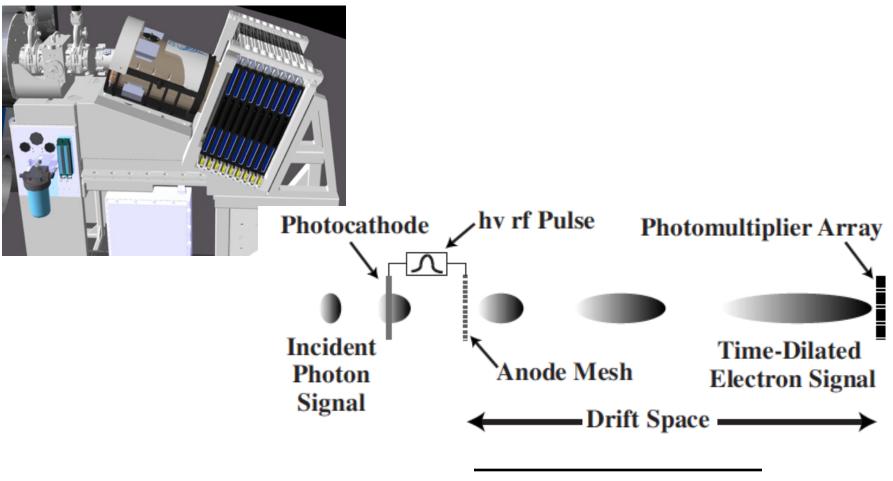
Proton radiography: magnetic fields around the magnetized jet





NIF's gated x-ray framing camera DIXI has 10-ps temporal resolution and observes features not seen with NIF's last century gated x-ray cameras



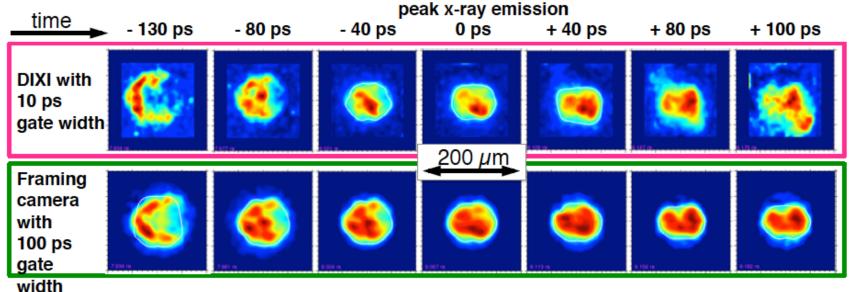


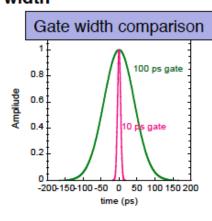
^{*}T. J. Hilsabeck et al., RSI 81, 10E317 (2010)

DIXI takes clearer pictures of the hot spot evolution around peak x-ray emission







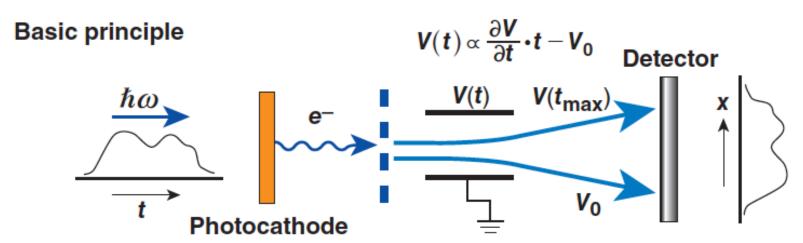


DIXI's 10X higher temporal resolution (reduced temporal blur) reveals details in the evolution of implosions at NIF never before possible, using the slower cameras.

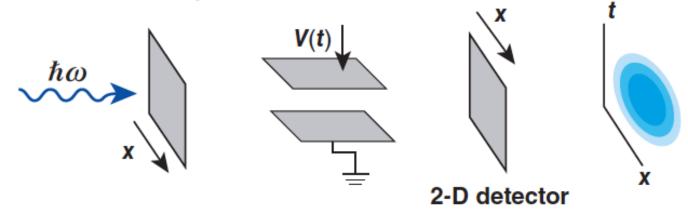
A streak camera provides temporal resolution of 1-D data







A streak camera can provide 2-D information



Outline



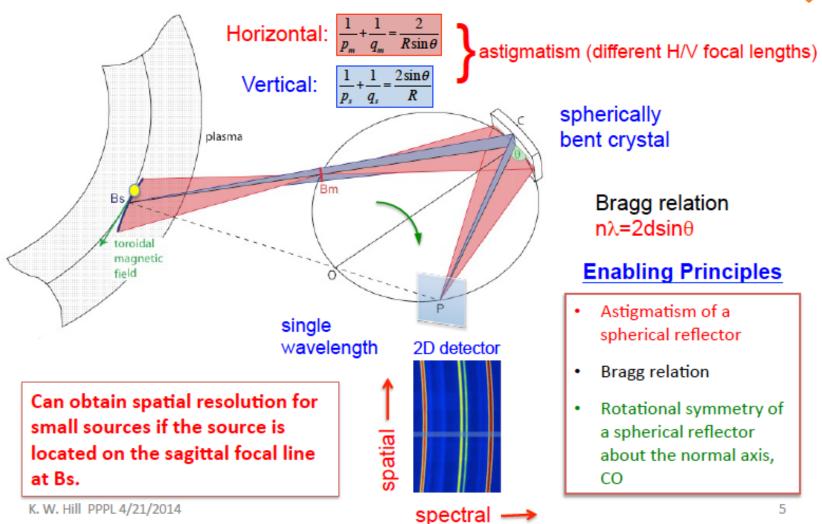


- Basic diagnostic building blocks
 - Electromagnetic field
 - Proton radiography
 - Particles
 - RCF stack / Proton activation pack / Electron spectrometer / Thomson parabola
 - X rays
 - Pinhole camera (time integrated) / 1-D streak camera / 2-D framing camera
 - Plasma conditions
 - High-resolution x-ray spectroscopy / Thomson scattering / Neutrons / x-ray radiography

High-resolution x-ray spectroscopy is well-established on Tokamaks for measuring plasma conditions







High-resolution x-ray spectroscopy is well-established on Tokamaks for measuring plasma conditions









X-ray Crystal Spectrometer Makes Debut at C-Mod

New Technique a Major Advance for ITER

PPPL/Alcator C-Mod collaboration has resulted in the demonstration of a greatly improved X-ray crystal spectrometer for application to ITER and fusion reactors. Experiments conducted by a PPPL/MIT team in April mark the beginning of a new era in the ability of such devices to determine radial profiles of the ion temperature and the rotational velocity of high temperature plasmas without the need for diagnostic beams. Their success



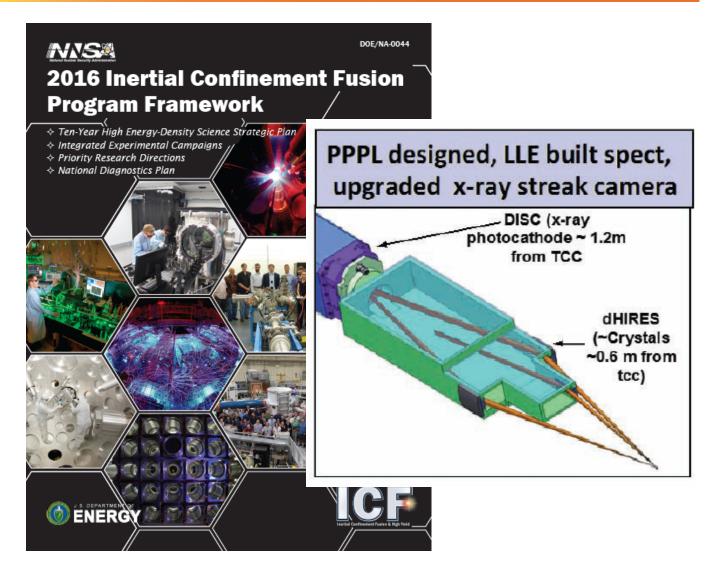
From the left are: Alex Ince-Cushman, MIT; Ken Hill, PPPL; Manfred Bitter, PPPL; John Rice, MIT; and Christian Broennimann of the Paul Scherrer Institute in Switzerland.

will benefit substantially ITER and other advanced fusion energy systems.

impurity by the pattern of frequencies, or spectrum, of the light emitted and they can determine the

PPPL designed high-resolution x-ray spectrometer has been identified as one of the 8 National Transformative Diagnostics

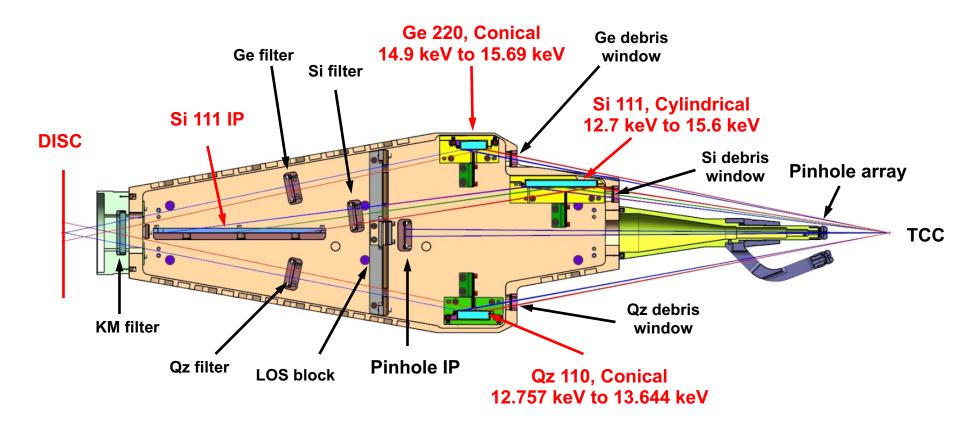




DHIRES contains three spectrometer channels and one imaging channel





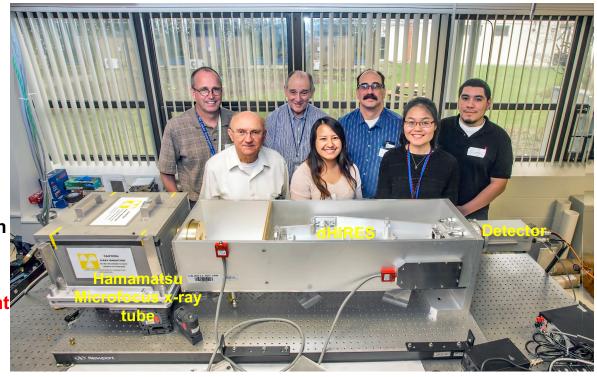


DHIRES is fully calibrated at the PPPL x-ray lab*





- Source Alignment
- Crystal Evaluation
- Energy Calibration and Crystal Dispersion
- Spectrum Manipulation
- Source Displacement and Insertion
 Error
- Absolute Throughput Measurement and NIF Signal Level Prediction
- Optical Alignment



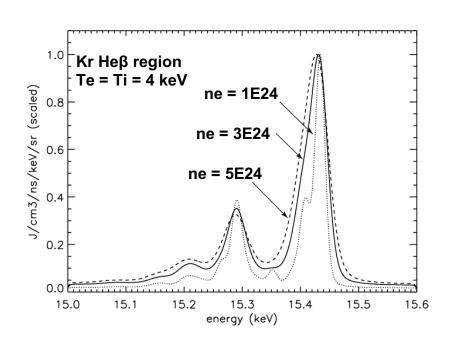
The spectral resolution is ~10 eV

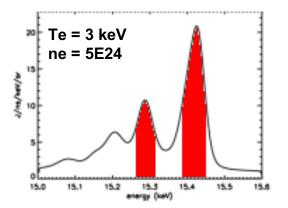
Stark broadening of Kr He β measures $n_{\rm e}$ and ratio of Helike resonance line to Li-like satellite provides $T_{\rm e}$

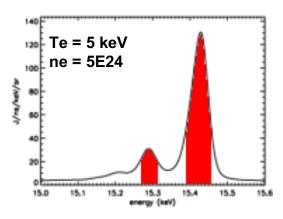




SCRAM calculation for a SymCap implosion Cr: D. Liedahl



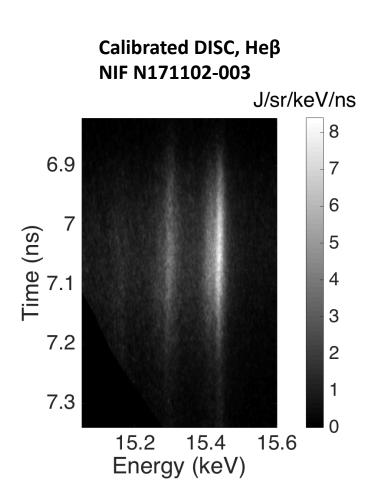


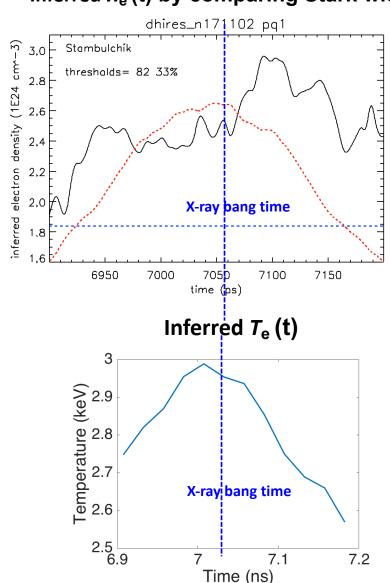


Stark broadening of Kr He β measures $n_{\rm e}$ and ratio of Helike resonance line to Li-like satellite provides $T_{\rm e}$







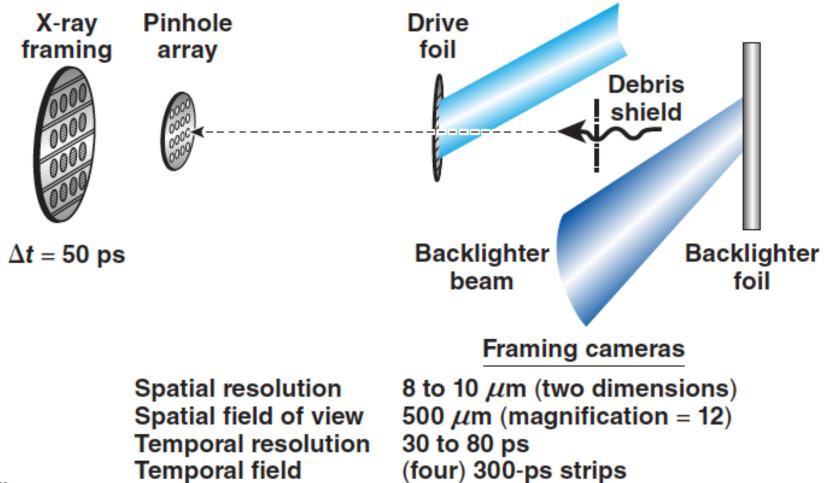


9

X-ray radiography, coupled with x-ray framing camera, is used to radiograph instability growth



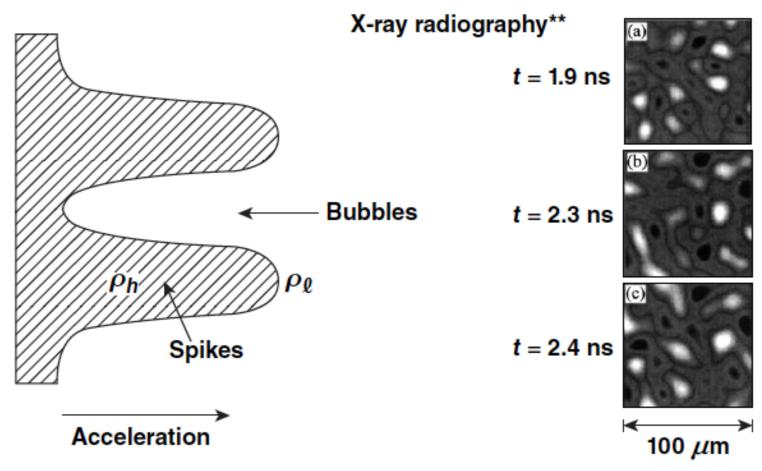
 The backlighter transmission depends on wavelength and target optical depth; mass perturbations lead to optical-depth perturbations





A bubble merger is predicted in the nonlinear phase of the RT instability*





X-ray photons are sensitive to density modulations.

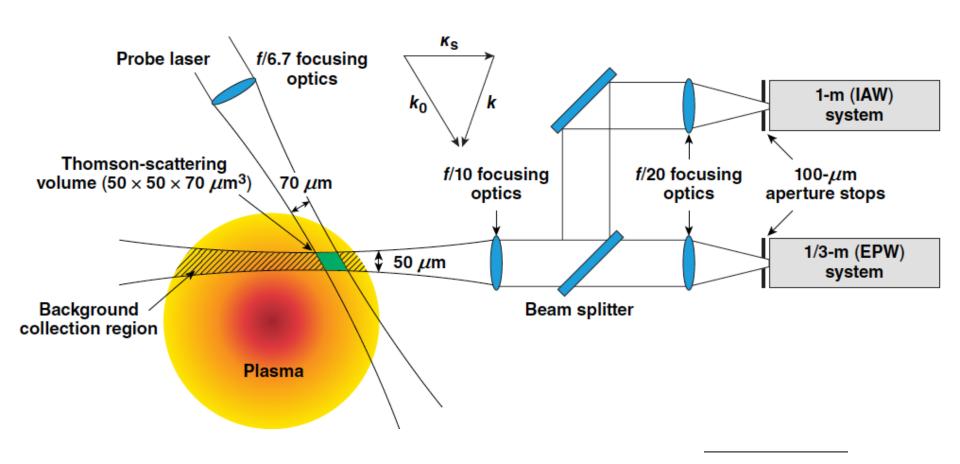
^{*}D. Oron et al., Phys. Plasmas 8, 2883 (2001);

U. Alon et al., Phys. Rev. Lett. 72, 2867 (1994).

^{**}V. A. Smalyuk et al., Phys. Rev. Lett. 81, 5342 (1998).

Optical Thomson scattering is used to diagnose local plasma conditions by observing the spectrum of light scattered from a probe beam over a small interaction volume



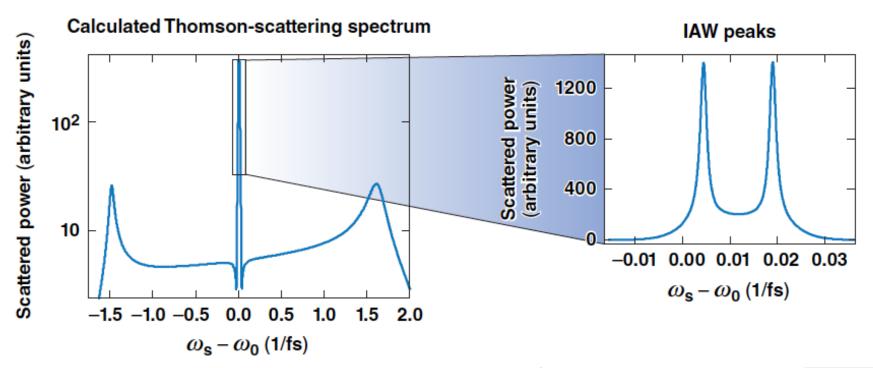


*R. Follet et al, RSI 87, 11E401 (2016)

Plasma parameters are inferred by comparing measured Thomson-scattering spectra to calculated spectra







Approximate frequency shift for scattering from

EPW's:
$$\Delta\omega_{\pm} = (\omega_{s} - \omega_{0}) = \pm\sqrt{\omega_{pe}^{2} + 3k^{2} v_{te}^{2}}$$

IAW's: $\Delta\omega_{\pm} = (\omega_{s} - \omega_{0}) = \pm kc_{s} - \vec{k} \cdot \vec{u}$

Electron density and temperature, ion temperature, and flow velocity were measured inside the magnetized jet





OMEGA beams:

Ring 1: 17, 22, 46, 56, 61

Ring 2: 11, 26, 31, 55, 68

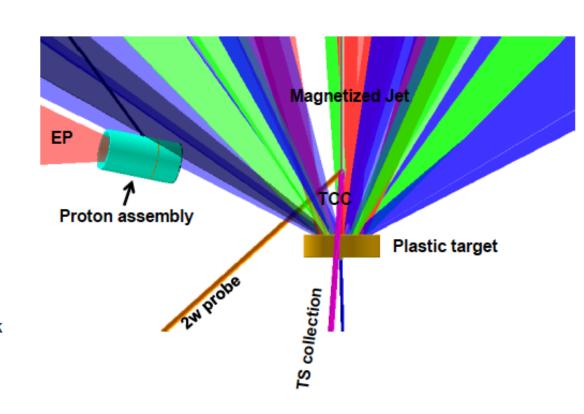
• Ring 3: 10, 13, 20, 28, 33, 48, 52, 58, 65, 66

OMEGA EP BL:

- Maximum energy, 10 ps (may switch to 1 ps)
- · Best focus
- 8 mm from TCC toward H7

Diagnostics:

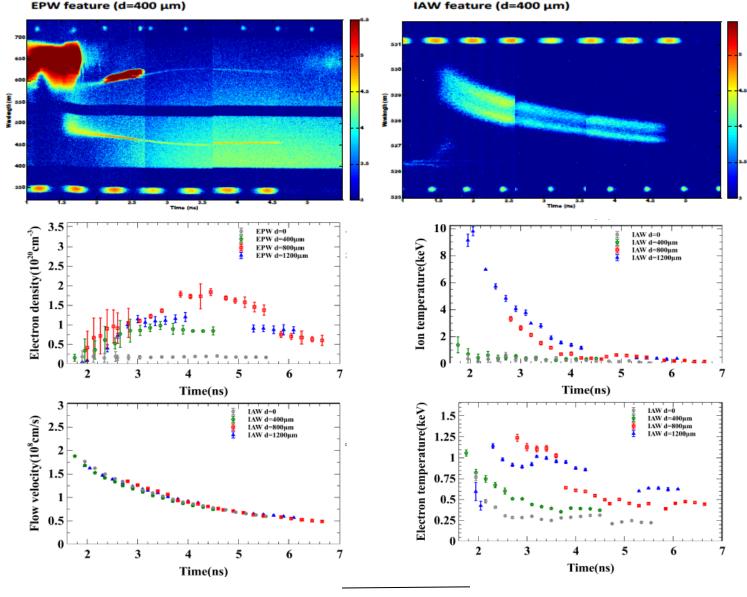
- TIM 5: Titled CPRM as the proton detector Distance is 16.5 cm from TCC
- TIM 6: TS collection system
- TIM 3: TTPS for the main plastic target
- TIM 4: XRFC (changed from TIM 2 to TIM 4)
- H2: EP proton generation assembly target stalk



9

Electron density and temperature, ion temperature, and flow velocity were measured inside the magnetized jet



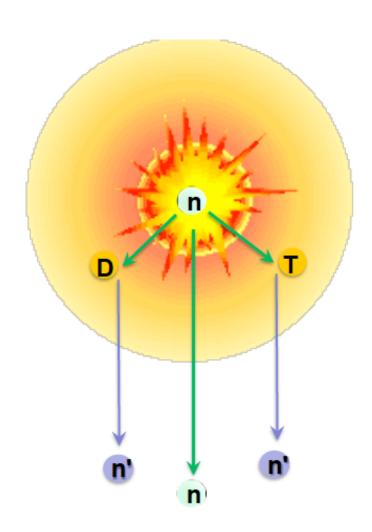


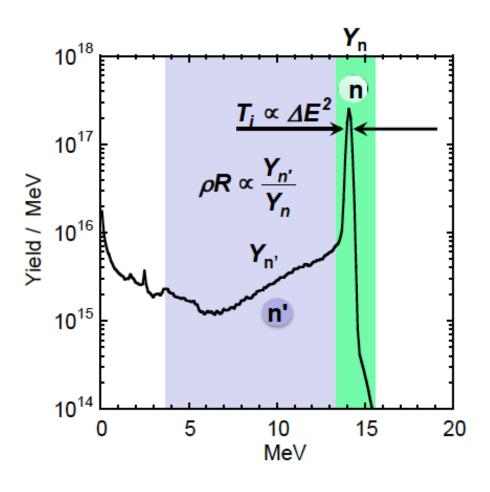
*L. Gao et al., The Astrophysical Journal Letters, 873:L11, 2019



The neutron spectrum provides information on ρR , Ti and yield – essential for assessing the implosion performance







Summary/Conclusions

A large suite of diagnostics have been developed for HEDP facilities



- HEDP systems generate some or all of
 - Visible light
 - UV and x-ray photons
 - Charged particles
 - Neutrons
 - Strong fields
- A comprehensive diagnostic suite makes it possible to learn a great deal about the systems: field strength and their impact, plasma parameters (n_e, n_i, T_e, v), particles, instabilities, yield, etc...
- Diagnosing HED systems require very high temporal (sub-ns, ps)
 and spatial (~10 μm) resolution
- Advances in technology and diagnostics enable understanding of new physics

